

**DARK FORCES  
AND DISCOVERY OPPORTUNITIES  
AT INTENSE ELECTRON BEAMS**

**NATALIA TORO  
PERIMETER INSTITUTE**

**IEB 2015  
CORNELL UNIVERSITY**



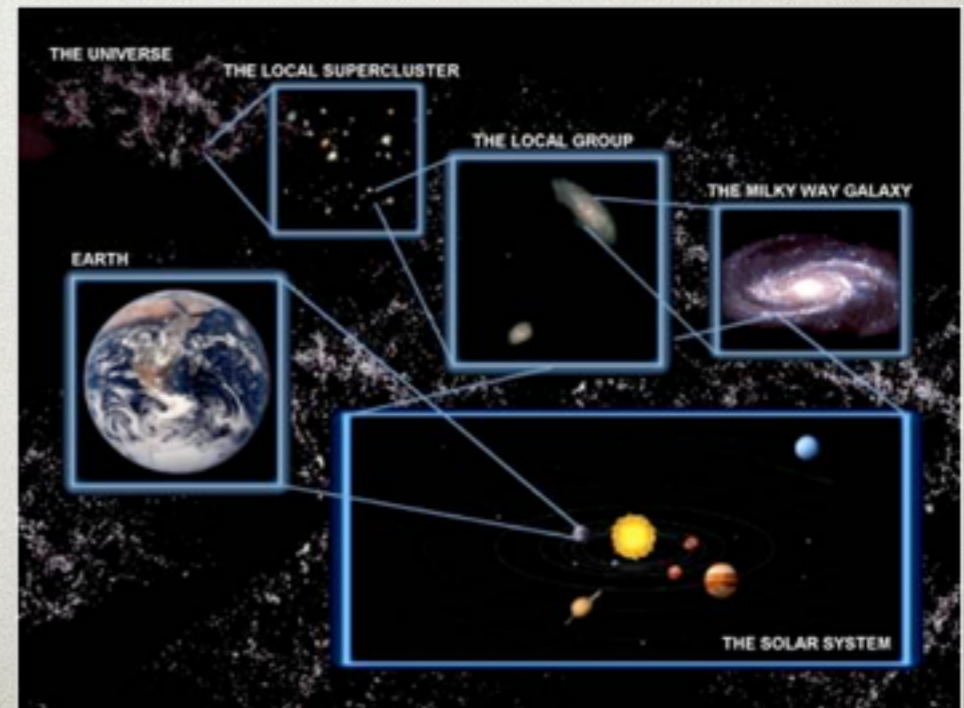
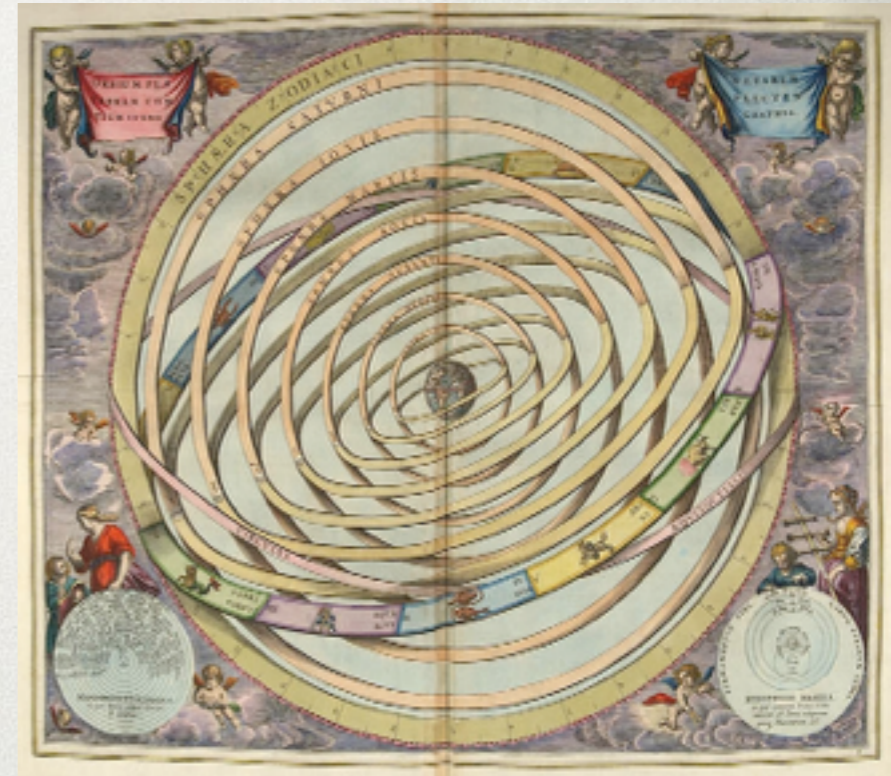
# COPERNICAN PARTICLE PHYSICS?

extension of Standard Model?

(superpartner, new weak multiplet...)



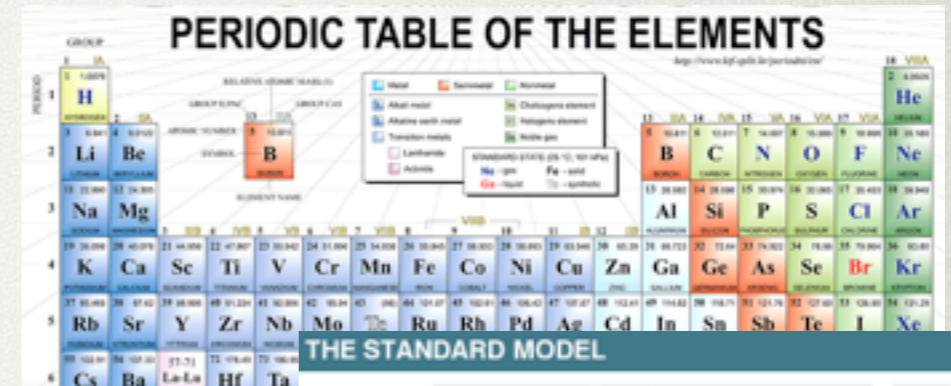
Completely new physics?





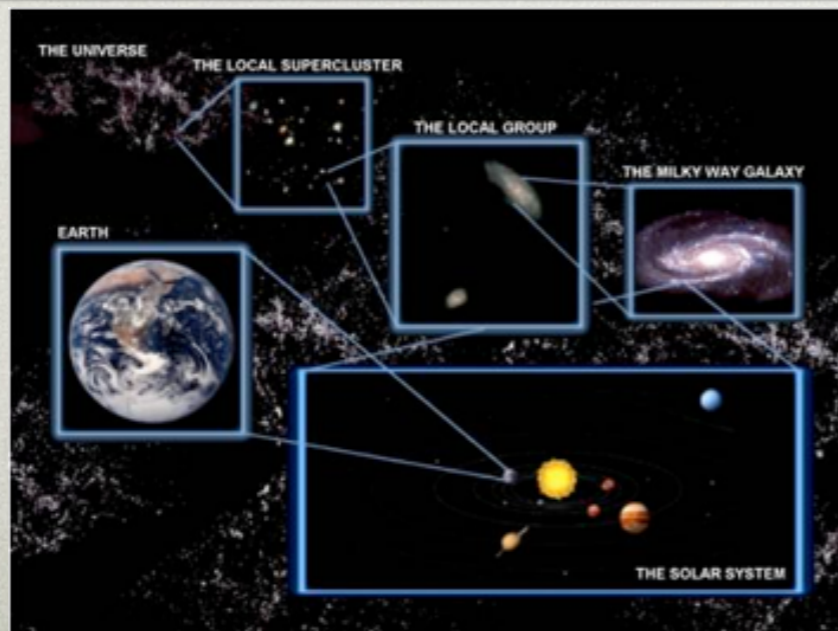
# COPERNICAN PARTICLE PHYSICS?

extension of Standard Model?  
(superpartner, new weak multiplet...)



Is there room for new physics not charged under Standard Model forces?  
What do we know about such physics, and how can we learn more?

Bosons	
$\gamma$ photon	Force carriers
Z Z boson	
W W boson	
g gluon	



THE STANDARD MODEL													
Fermions			Bosons										
Quarks	<table border="1"> <tr> <td>u up</td> <td>c charm</td> <td>t top</td> </tr> <tr> <td>d down</td> <td>s strange</td> <td>b bottom</td> </tr> </table>	u up	c charm	t top	d down	s strange	b bottom	<table border="1"> <tr> <td><math>\gamma</math> photon</td> </tr> <tr> <td>Z Z boson</td> </tr> <tr> <td>W W boson</td> </tr> <tr> <td>g gluon</td> </tr> </table>	$\gamma$ photon	Z Z boson	W W boson	g gluon	?
u up	c charm	t top											
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$\gamma$ photon													
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Leptons	<table border="1"> <tr> <td><math>\nu_e</math> electron neutrino</td> <td><math>\nu_\mu</math> muon neutrino</td> <td><math>\nu_\tau</math> tau neutrino</td> </tr> <tr> <td>e electron</td> <td><math>\mu</math> muon</td> <td><math>\tau</math> tau</td> </tr> </table>	$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino	e electron	$\mu$ muon	$\tau$ tau	Higgs boson *Yet to be confirmed	...				
$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino											
e electron	$\mu$ muon	$\tau$ tau											
				?									



# SEARCHING FOR PHYSICS

## “OUTSIDE” THE STANDARD MODEL

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- Dark / Hidden Sectors: Mapping out the Possibilities
- Searching for Dark Forces
  - Multi-purpose collider experiments
  - Dedicated fixed-target experiments
- Light Dark Matter
  - An opportunity for very low-energy beams



# HOW TO LOOK FOR PHYSICS **FAR** BEYOND THE SM?

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Accessible **mass** isn't enough – need interactions

⇒ suppressed by high power of mass-scale at which interactions are generated

$$(\bar{\psi}_e \psi_e)_{SM} (\bar{\chi} \chi)_{new} / \Lambda^2$$

[analogous to approximate stability of proton in SM]

Even if  $\chi$  is light, large  $\Lambda \Rightarrow$  unobservable effect.

The **few operators with no  $\Lambda$ -suppression\*** present an opportunity to explore this physics

\*gauge-invariant combinations of SM and new fields with dimension  $< 4$



# THE “PORTALS”

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Neutrino Portal  $\epsilon_\nu (hL)\psi$  sterile neutrinos?

Higgs Portal  $\epsilon_h |h|^2 |\phi|^2$  exotic rare Higgs decays?

Vector Portal  
(kinetic mixing)  $\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$  [Holdom '86]

Generic low-energy remnants of *any* non-SM sector

*Only light-vector portal is truly accessible in low-energy production* (e & p couplings to h,  $\nu$  are small)



# AXIONS AND THE ALMOST-PORTALS

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A pseudo-scalar boson can have several interactions suppressed by only one power of mass

$$\text{Axion Portal} \left\{ \begin{array}{l} \frac{1}{\Lambda} G_{\mu\nu} \tilde{G}^{\mu\nu} a \\ \frac{1}{\Lambda} \bar{f} \gamma^\mu \gamma_5 f \partial_\mu a \\ \frac{1}{\Lambda} F_{\mu\nu} \tilde{F}^{\mu\nu} a \end{array} \right.$$

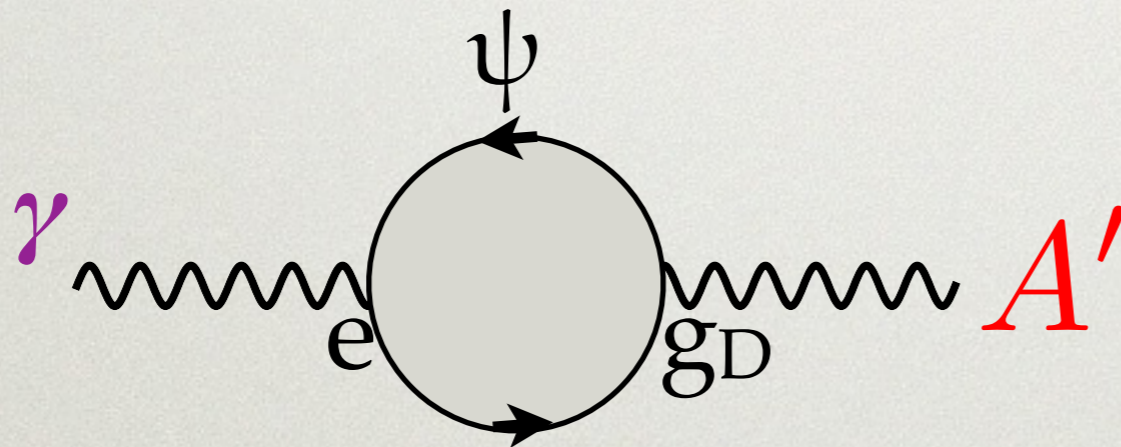
Even for large  $\Lambda$ , coherent-field effects can compensate for weak coupling



# SOURCES AND SIZES OF

## KINETIC MIXING $\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$

- If absent from fundamental theory, can still be generated by **perturbative** (or non-perturbative) quantum effects
  - Simplest case: one heavy particle  $\psi$  with both **EM charge** & **dark charge**



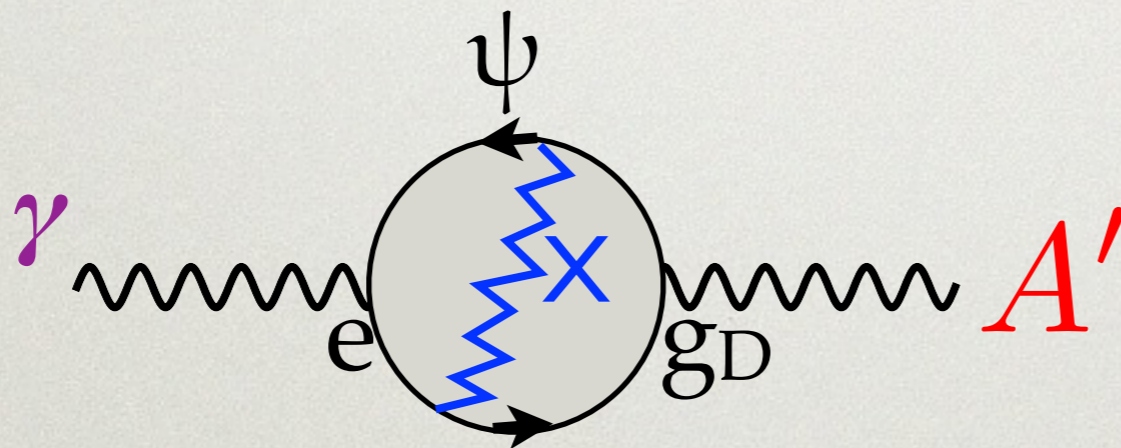
generates  $\epsilon \sim \frac{e g_D}{16\pi^2} \log \frac{m_\psi}{M_*} \sim 10^{-2} - 10^{-4}$



# SOURCES AND SIZES OF

## KINETIC MIXING $\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$

- If absent from fundamental theory, can still be generated by **perturbative** (or non-perturbative) quantum effects
  - In Grand Unified Theory, symmetry forbids tree-level & 1-loop mechanisms. **GUT-breaking** enters at 2 loops



generating  $\epsilon \sim 10^{-3} - 10^{-5}$

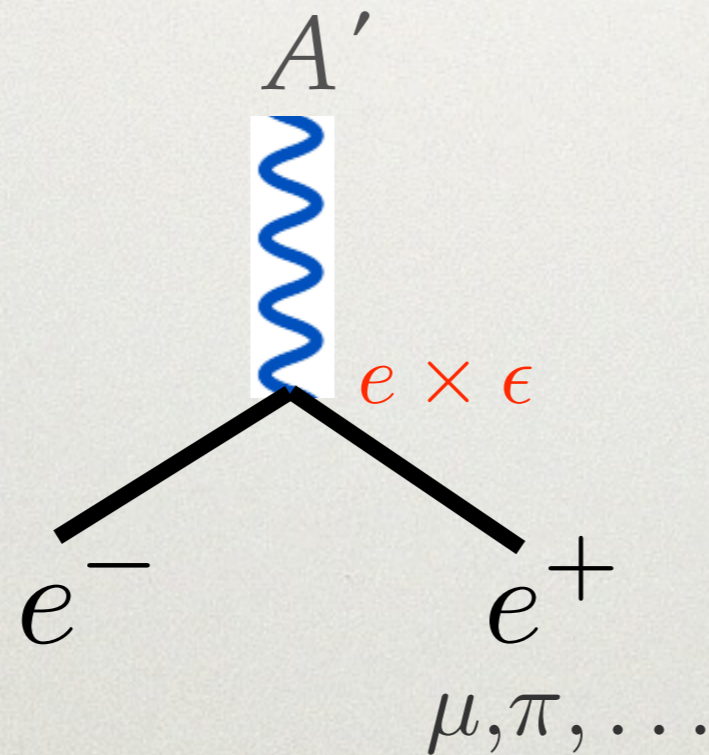
( $\rightarrow 10^{-7}$  if both  $U(1)$ 's are in unified groups)



# EFFECTS OF KINETIC MIXING

$$\frac{1}{2} \epsilon_Y F_{\mu\nu}^Y F'^{\mu\nu}$$

Regardless of where it comes from, kinetic mixing can always be re-interpreted as (mainly) giving matter of electric charge  $qe$  an  $A'$  coupling  $\propto q\epsilon e$



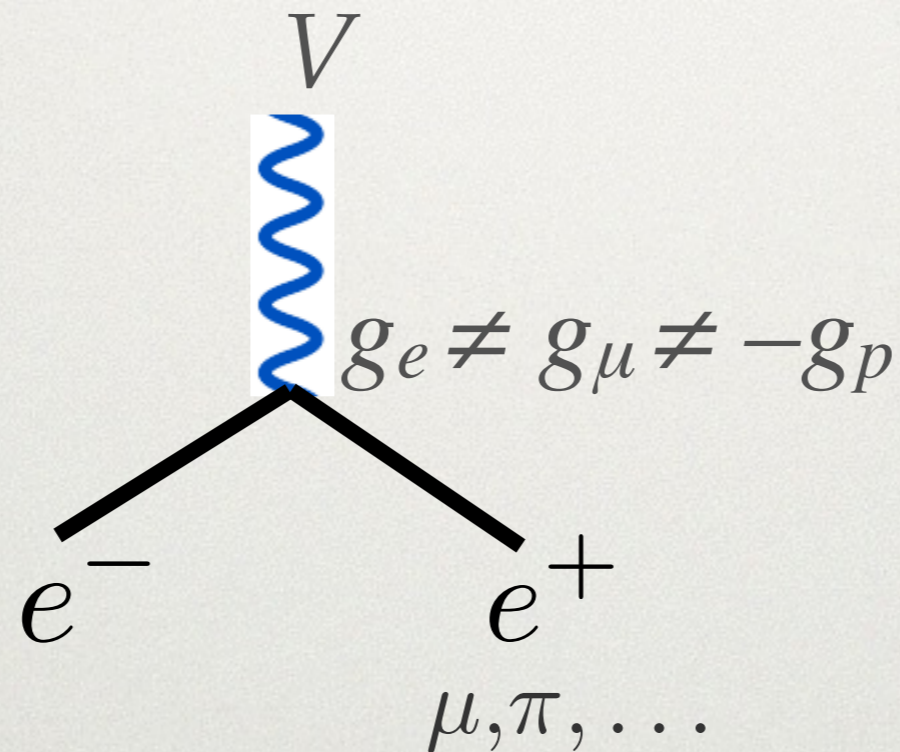
Dark matter can have an independent coupling  $g_D$  to  $A'$



# BYCATCH

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New, weak gauge forces of the Standard Model (e.g. to B-L) can be found by the same kinds of searches.



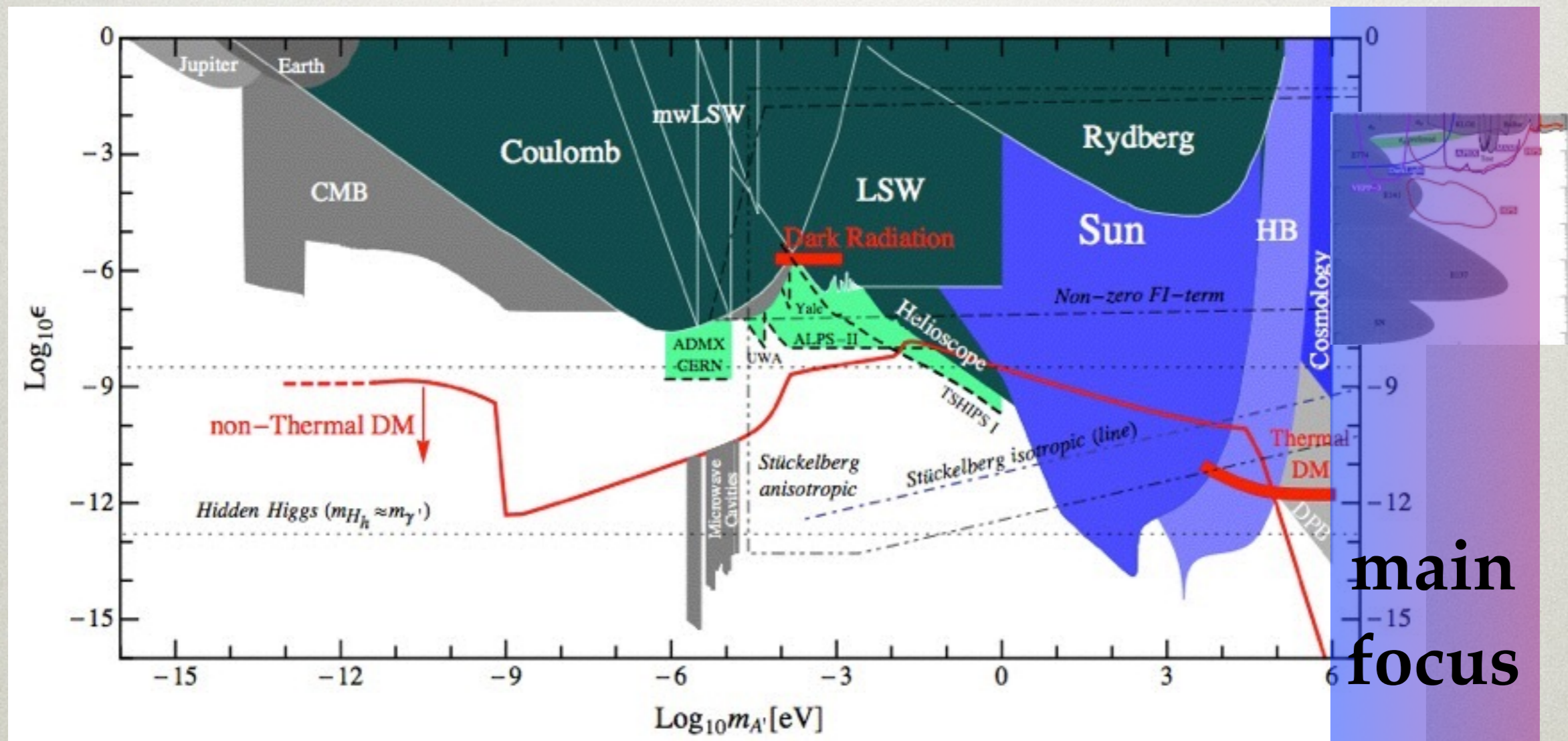
Important to keep this in mind in comparing e.g.  $p$  to  $e^\pm$  beams

May be relevant to e.g. muonic hydrogen anomaly



# WIDE PARAMETER SPACE: HIDDEN VECTORS

Nuclear  
scale



[Figure from 2013 Intensity Frontier report – Javier Redondo]

new particles



# SOURCES AND SIZES OF MASS TERM

---

- MeV-to-GeV is **allowed** at couplings  $>10^{-7}$
- Possible origin: related to  $M_Z$  by small parameter
  - e.g. supersymmetry+kinetic mixing  $\Rightarrow$  scalar coupling to SM Higgs, giving

$$m_{A'} \sim \sqrt{\epsilon} M_Z \lesssim 1\text{GeV}$$

[e.g. Cheung, Ruderman, Wang, Yavin; Katz, Sundrum; Morrissey, Poland, Zurek]

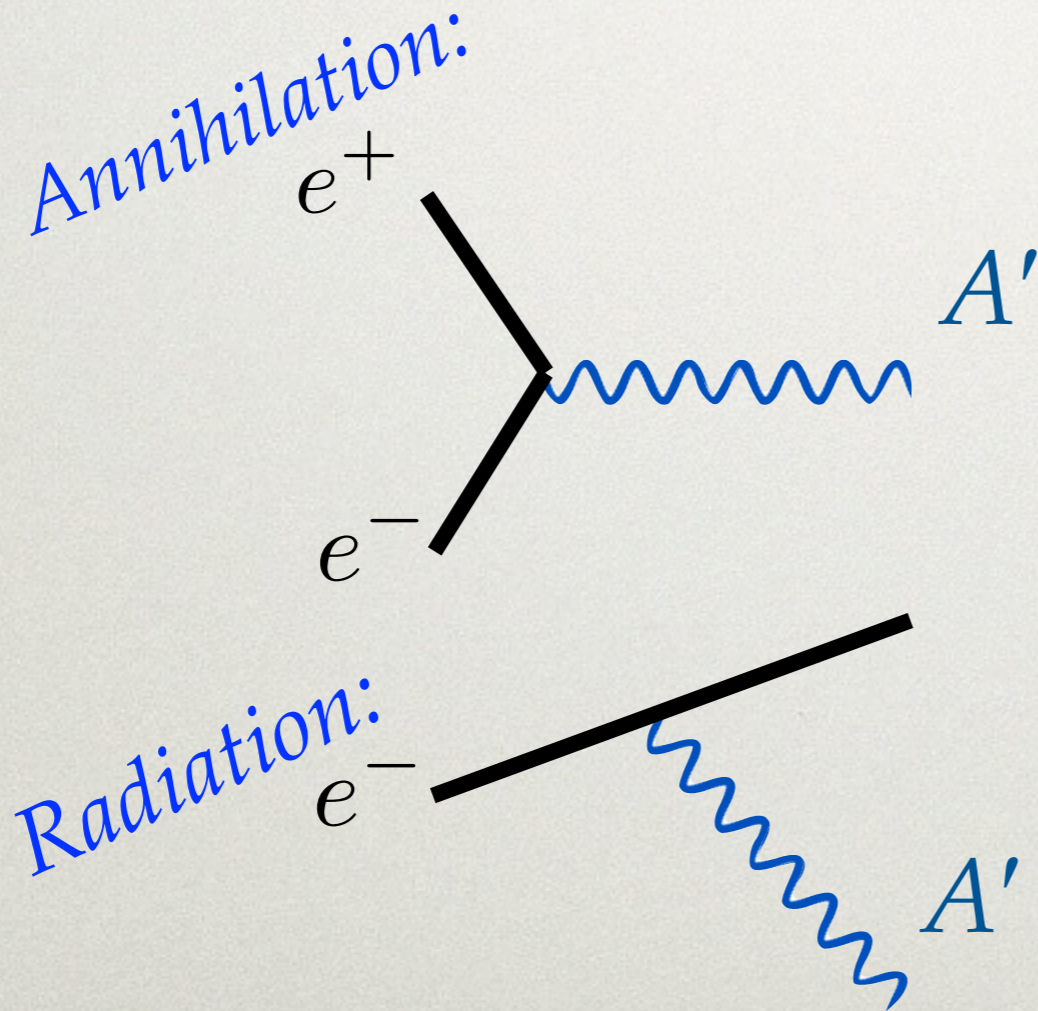
- motivated by  $g-2$  and dark matter anomalies
- A particularly **relevant** and **accessible** range to explore



# A FIELD GUIDE TO DARK FORCES PRODUCTION

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(like ordinary radiation of light, but suppressed by  $\epsilon$ )



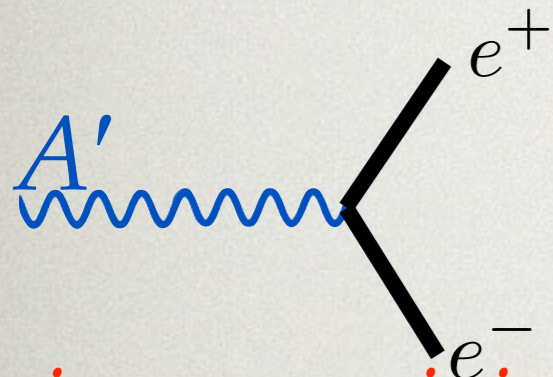
“Wherever there’s a photon,  
there’s a dark photon”



# A FIELD GUIDE TO DARK FORCES

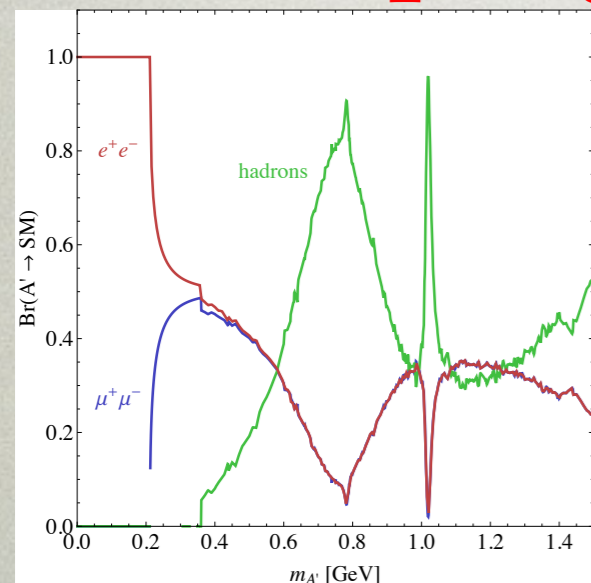
## DECAY

“Minimal” Decay:

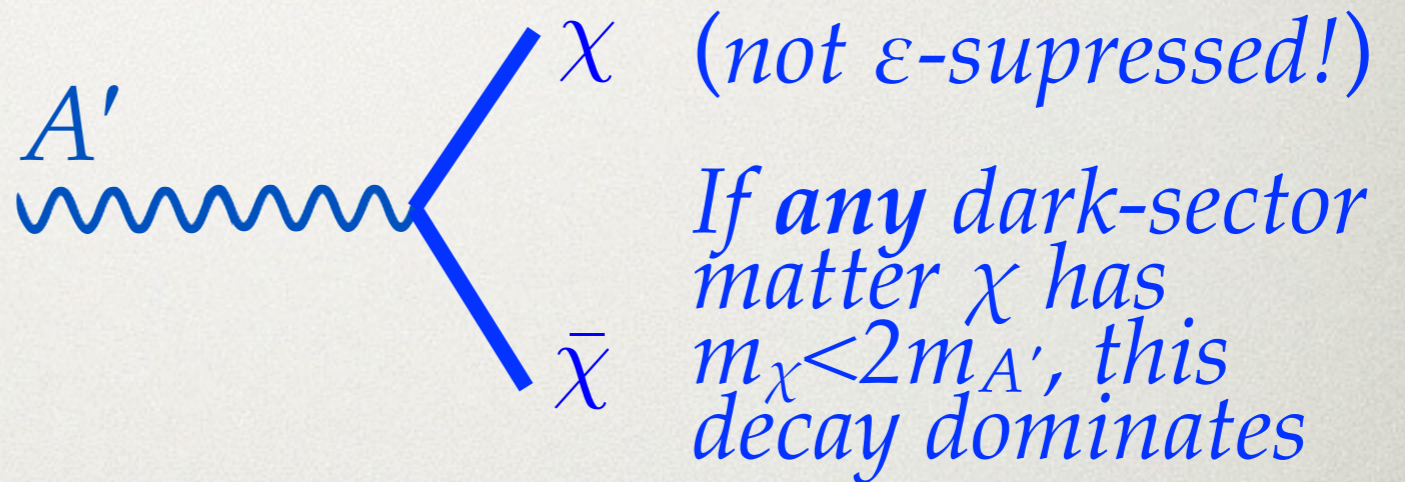


*via same mixing operator as production  
 $\Rightarrow$  tiny width*

$$\Gamma \sim \epsilon^2 \alpha m_{A'}$$



“Generic” Decay:



Two cases:

- $\chi$  stable & invisible

- $\chi$  decays into SM particles,  $A' \rightarrow >2$  charged particles

searches at BaBar and KLOE

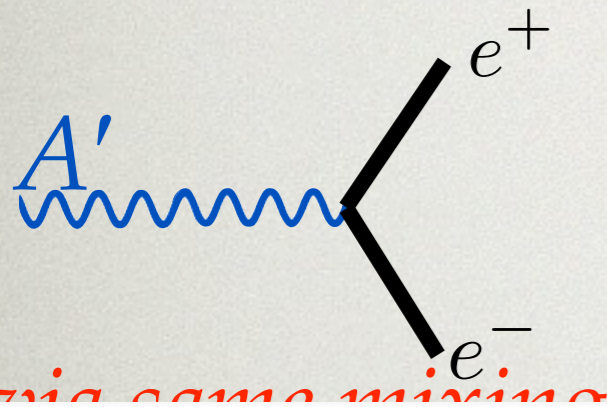
To test “dark sector” idea, we need to search for both!



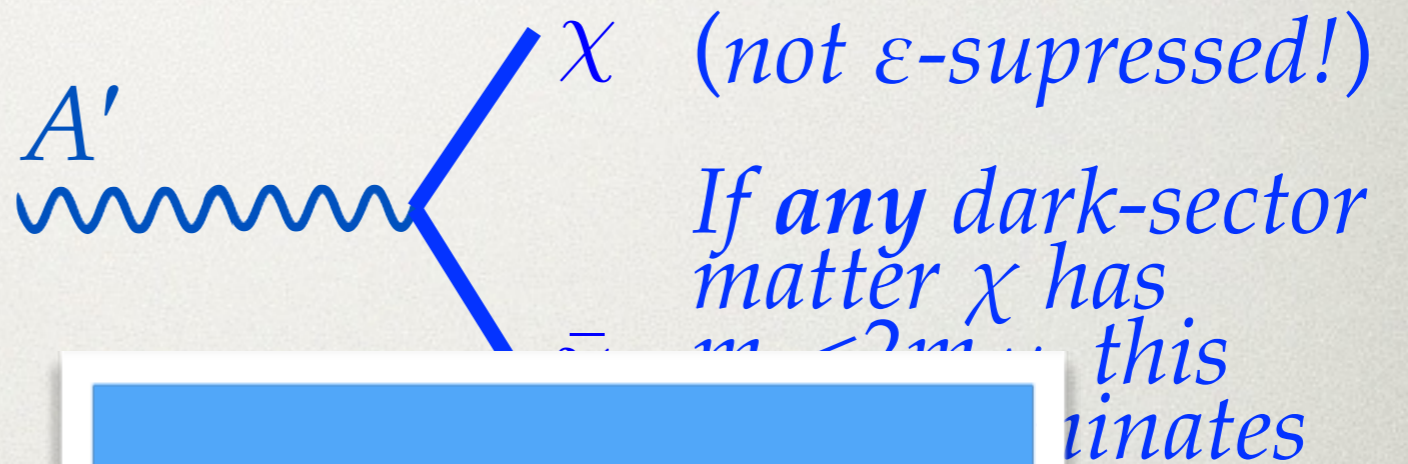
# A FIELD GUIDE TO DARK FORCES

## DECAY

“Minimal” Decay:

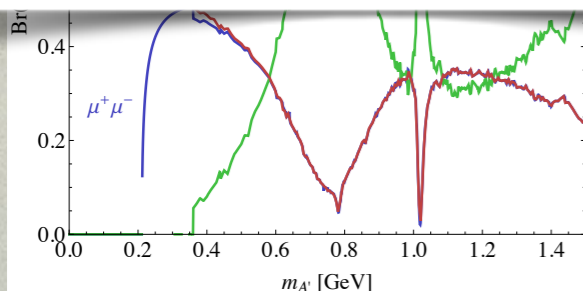


“Generic” Decay:



Major advances in last 5 years!

Huge upcoming opportunity!



$T_W$

$A \rightarrow \gamma\gamma$  charged particles searches at BaBar and KLOE

To test “dark sector” idea, we need to search for both!

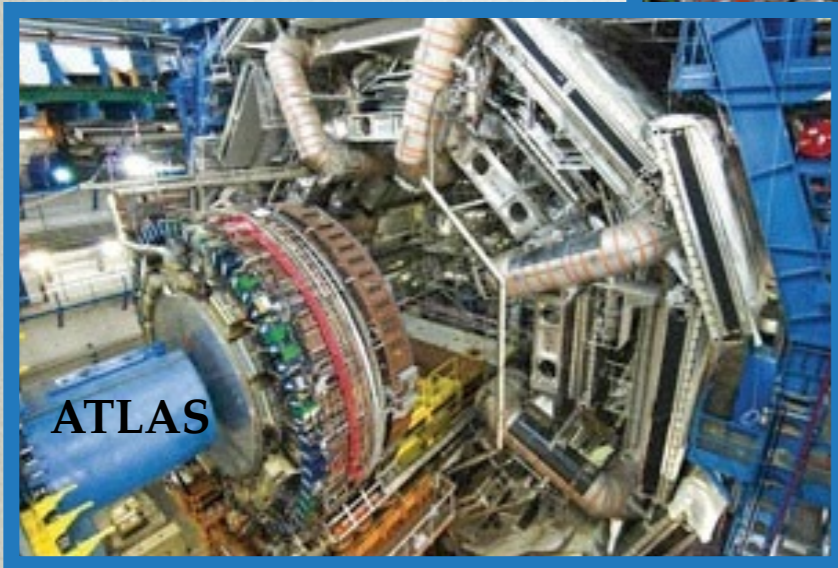


# AN EXPERIMENTAL RENAISSANCE

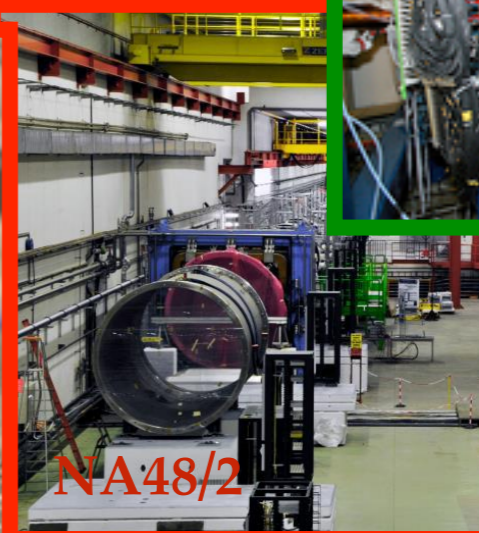
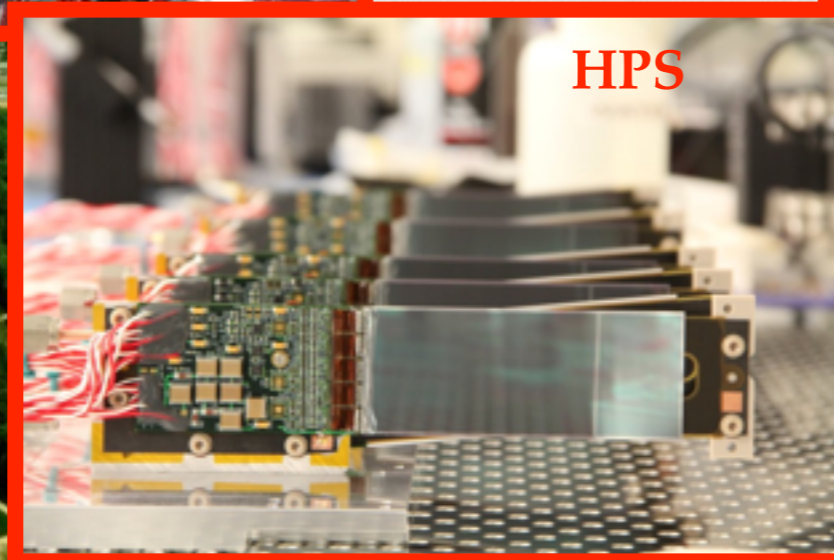
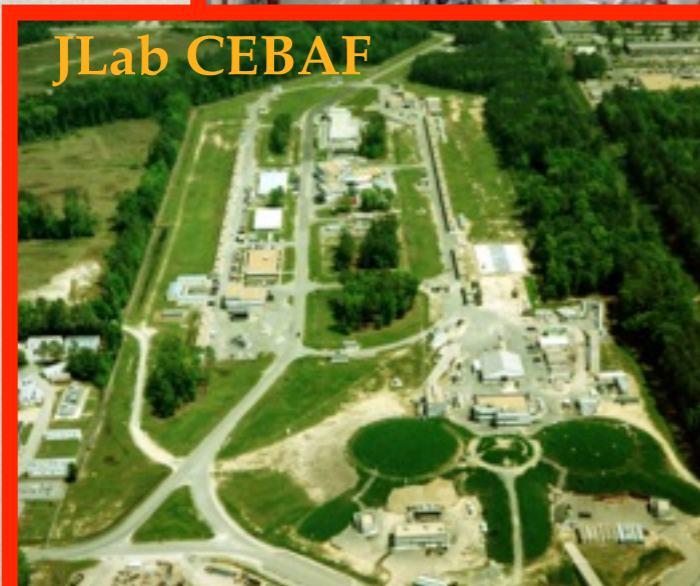
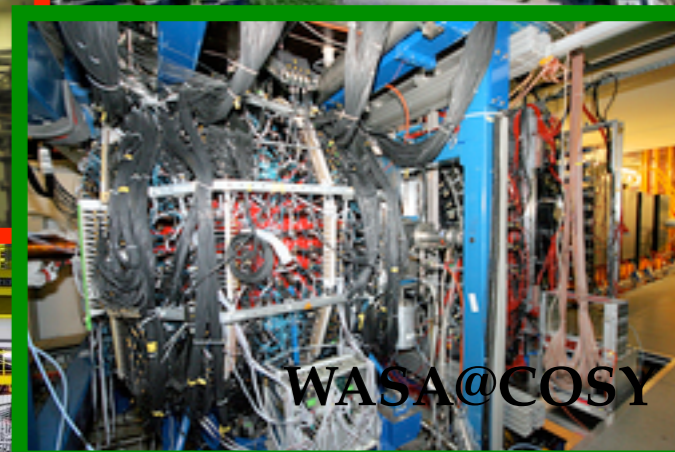
High-energy  
colliders



High intensity  
colliders



Fixed  
Target

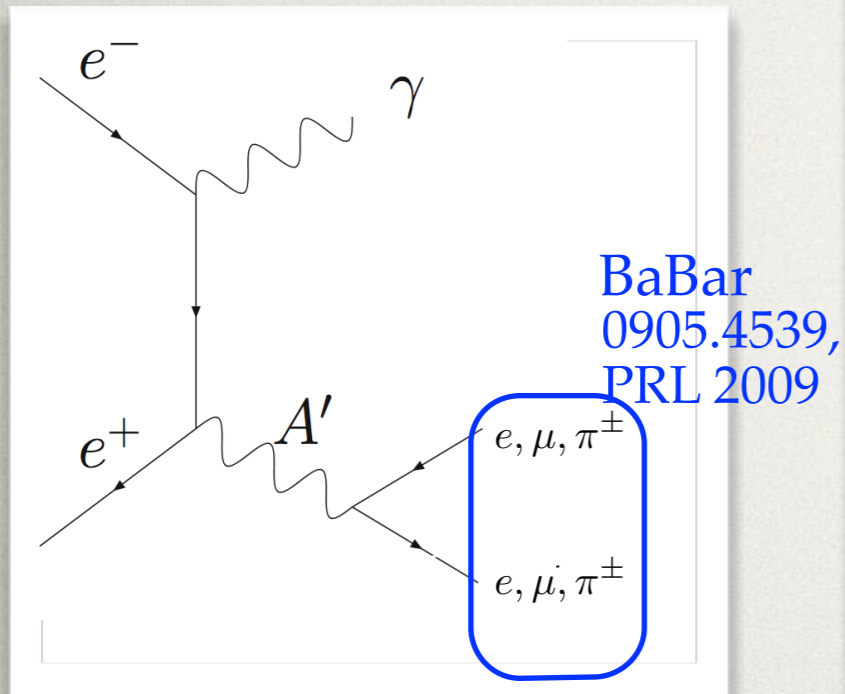




# COLLIDER PRODUCTION

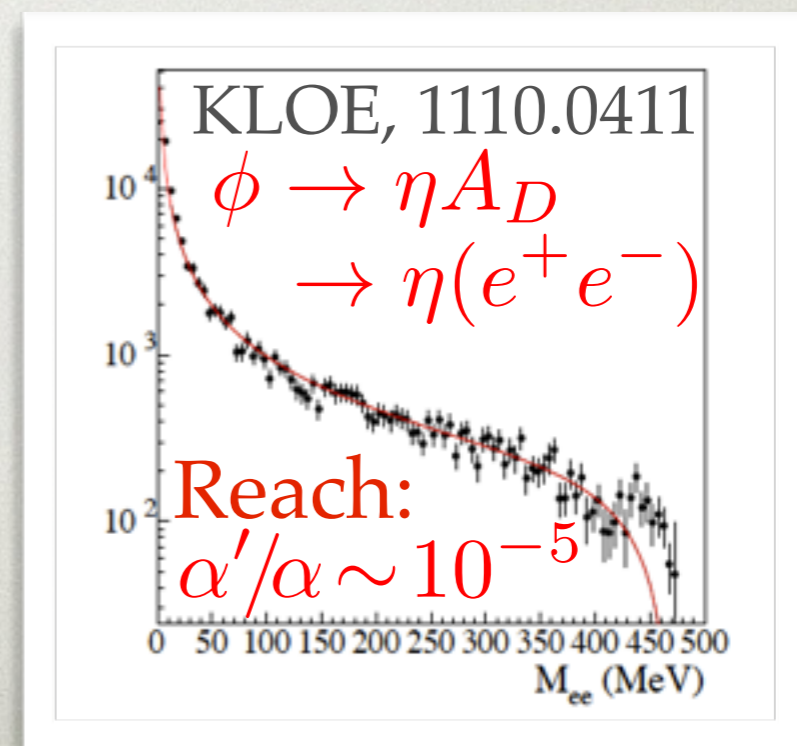
Radiative return

Rare meson decays



$X \rightarrow YU$	$n_X$	$m_X - m_Y$ (MeV)	$\text{BR}(X \rightarrow Y + \gamma)$	$\text{BR}(X \rightarrow Y + \ell^+\ell^-)$	$\epsilon \leq$
$\eta \rightarrow \gamma U$	$n_\eta \sim 10^7$	547	$2 \times 39.8\%$	$6 \times 10^{-4}$	$2 \times 10^{-3}$
$\omega \rightarrow \pi^0 U$	$n_\omega \sim 10^7$	648	8.9%	$7.7 \times 10^{-4}$	$5 \times 10^{-3}$
$\phi \rightarrow \eta U$	$n_\phi \sim 10^{10}$	472	1.3%	$1.15 \times 10^{-4}$	$1 \times 10^{-3}$
$K_L^0 \rightarrow \gamma U$	$n_{K_L^0} \sim 10^{11}$	497	$2 \times (5.5 \times 10^{-4})$	$9.5 \times 10^{-6}$	$2 \times 10^{-3}$
$K^+ \rightarrow \pi^+ U$	$n_{K^+} \sim 10^{10}$	354	-	$2.88 \times 10^{-7}$	$7 \times 10^{-3}$
$K^+ \rightarrow \mu^+ \nu U$	$n_{K^+} \sim 10^{10}$	392	$6.2 \times 10^{-3}$	$7 \times 10^{-8a}$	$2 \times 10^{-3}$
$K^+ \rightarrow e^+ \nu U$	$n_{K^+} \sim 10^{10}$	496	$1.5 \times 10^{-5}$	$2.5 \times 10^{-8}$	$7 \times 10^{-3}$

PLB706 (2012) 251-255

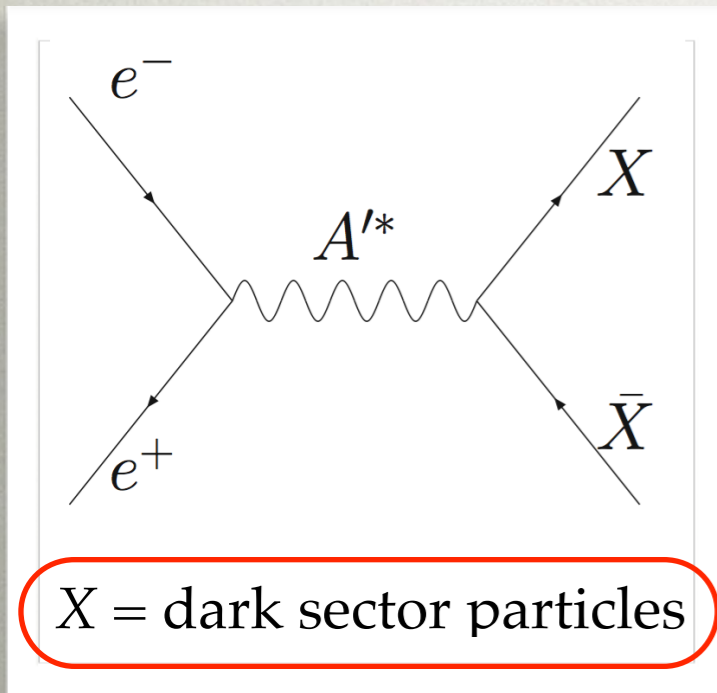




# WIDE BREADTH OF SEARCHES

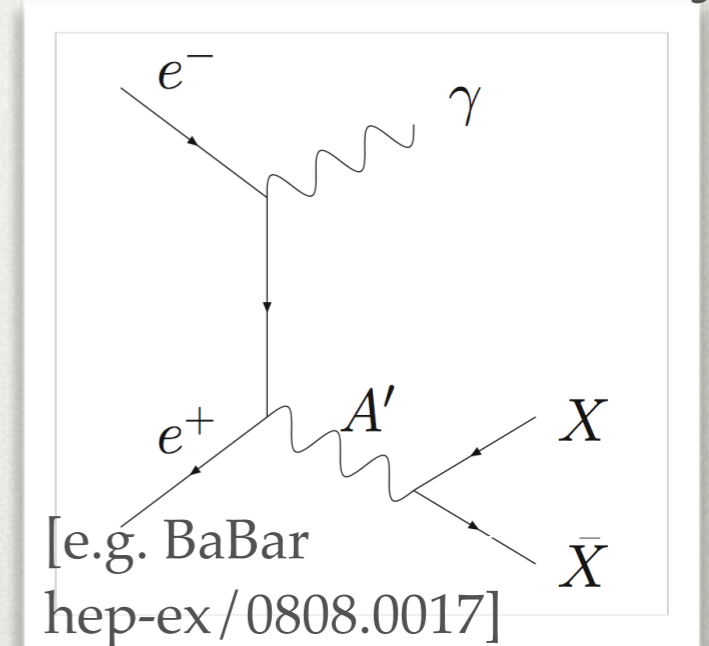
(just a few representative examples)

## Off-shell $A'$ portal

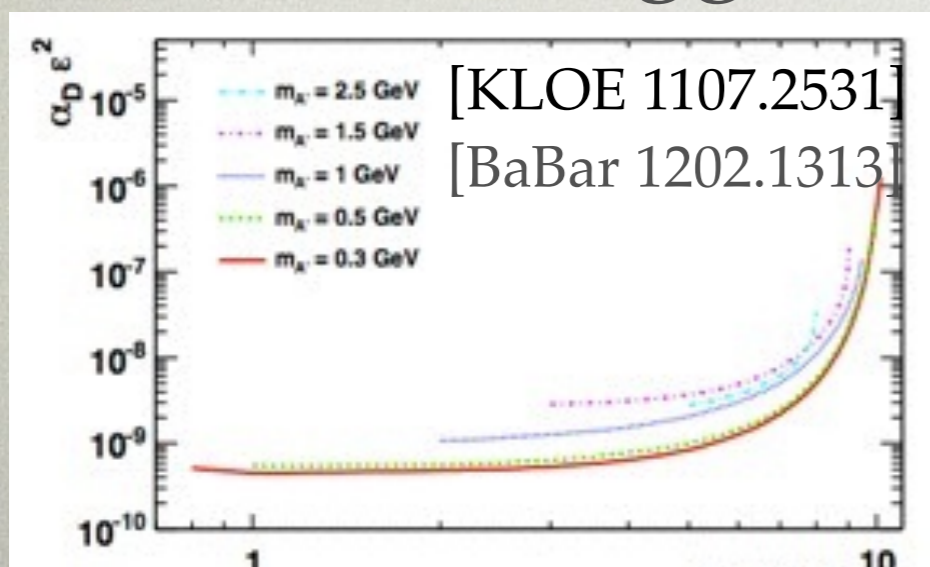


Potential to see rich hidden sectors in complex multi-body final states (searches ongoing at BaBar + several completed)

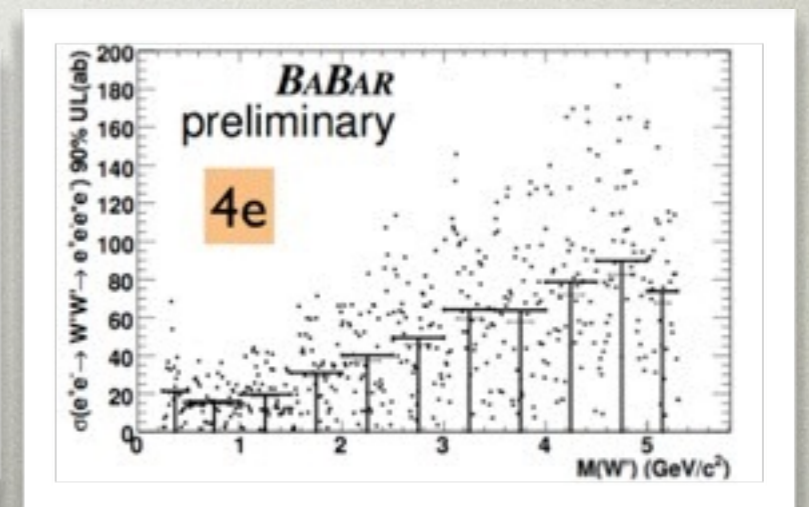
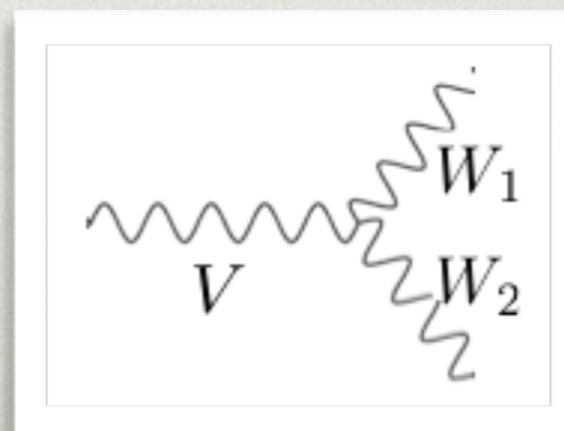
## Invisible Decay



## Vector + Higgs:



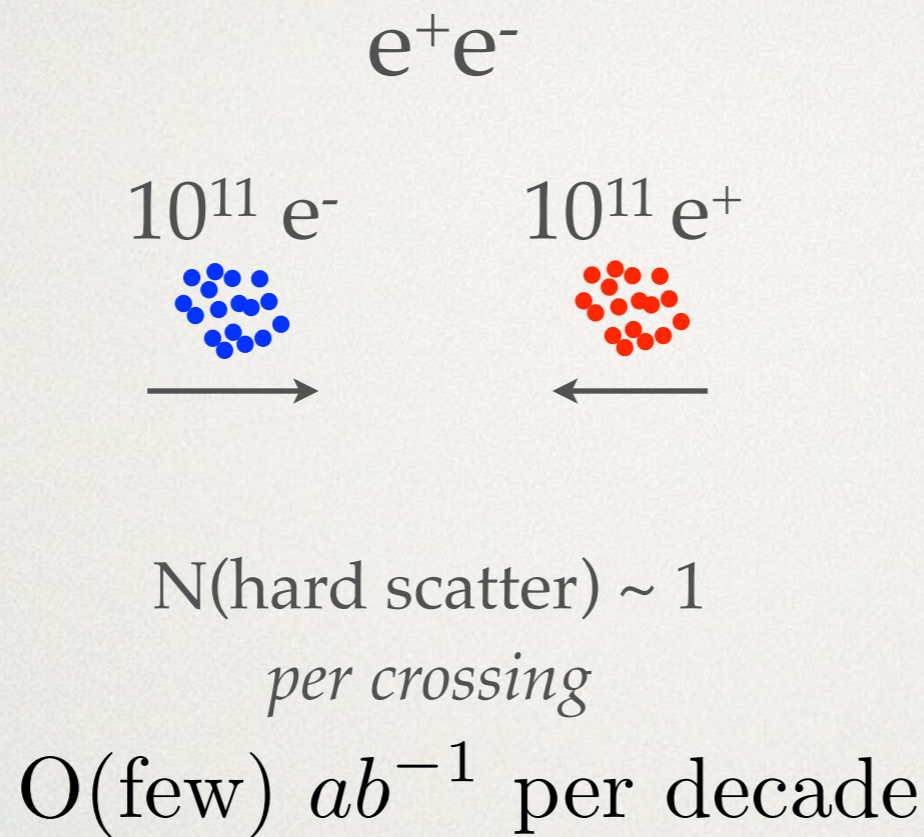
## Non-Abelian Dark Sector



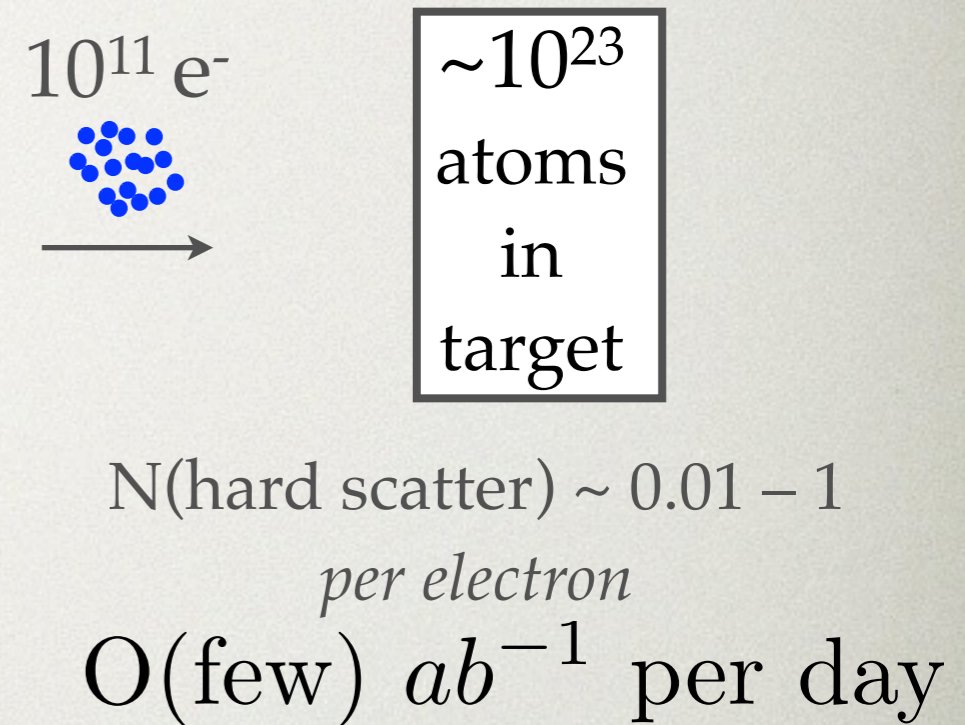


# GOING FURTHER: FIXED TARGET

## LUMINOSITY

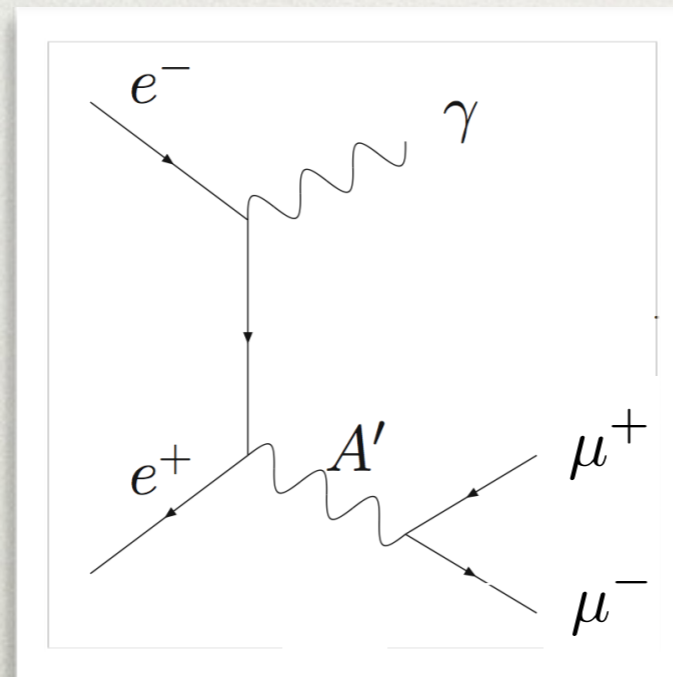


## Fixed-Target

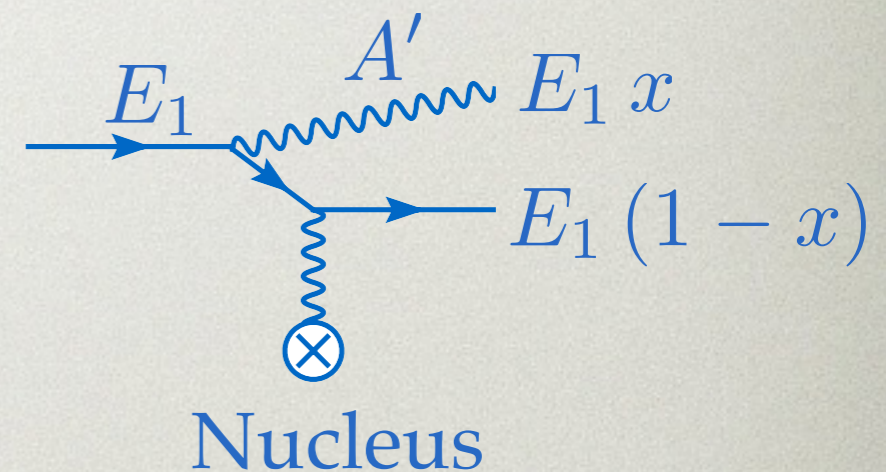


## CROSS-SECTION

- Scales as  $A'$  mass, not beam energy
- Coherent scattering from nucleus



$$\sigma \sim \frac{\alpha^2 \epsilon^2}{E^2} \sim O(10 \text{ fb})$$



$$\sigma \sim \frac{\alpha^3 Z^2 \epsilon^2}{m^2} \sim O(10 \text{ pb})$$



# Jefferson Lab Continuous Electron Beam Accelerator Facility

- Delivers beam up to 12 GeV to 4 experimental hall



Halls A,C up to  $100 \mu\text{A}$   
Hall B,D:  $1 \mu\text{A}$

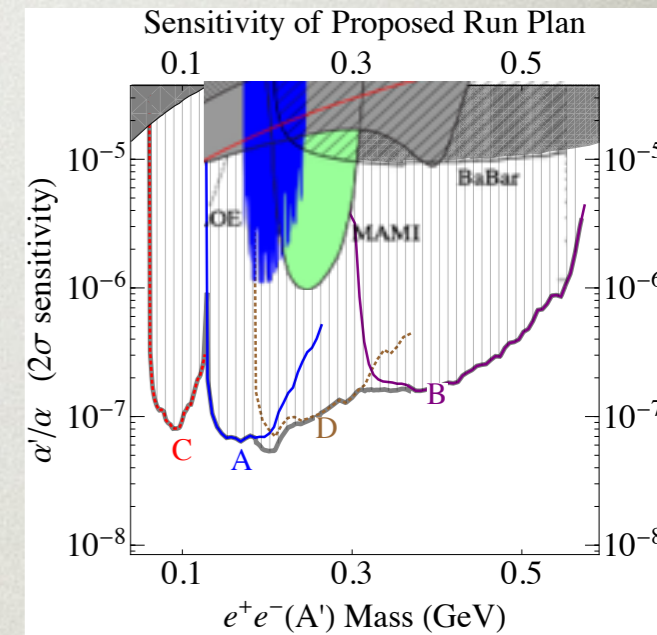
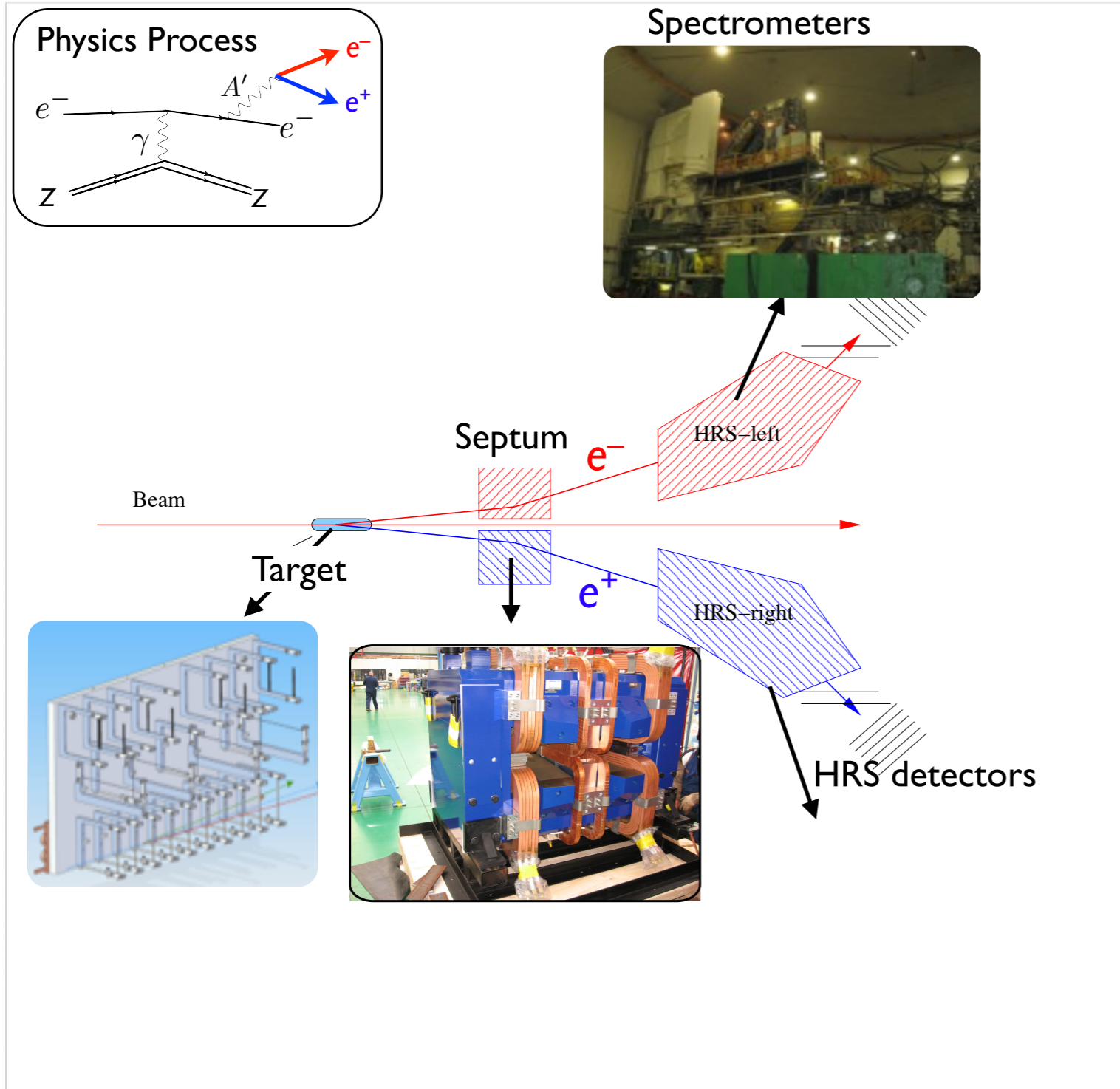
- 1.5 GHz RF  $\Rightarrow$  each hall gets bunch every 2–4 ns
- Commissioning beam to halls after recent energy upgrade & addition of Hall D
- Only multi-GeV continuous electron beam in the world!



# APEX

<http://hallaweb.jlab.org/experiment/APEX/>

Search for new gauge boson  $A'$  using Hall A high-resolution spectrometers (HRS)

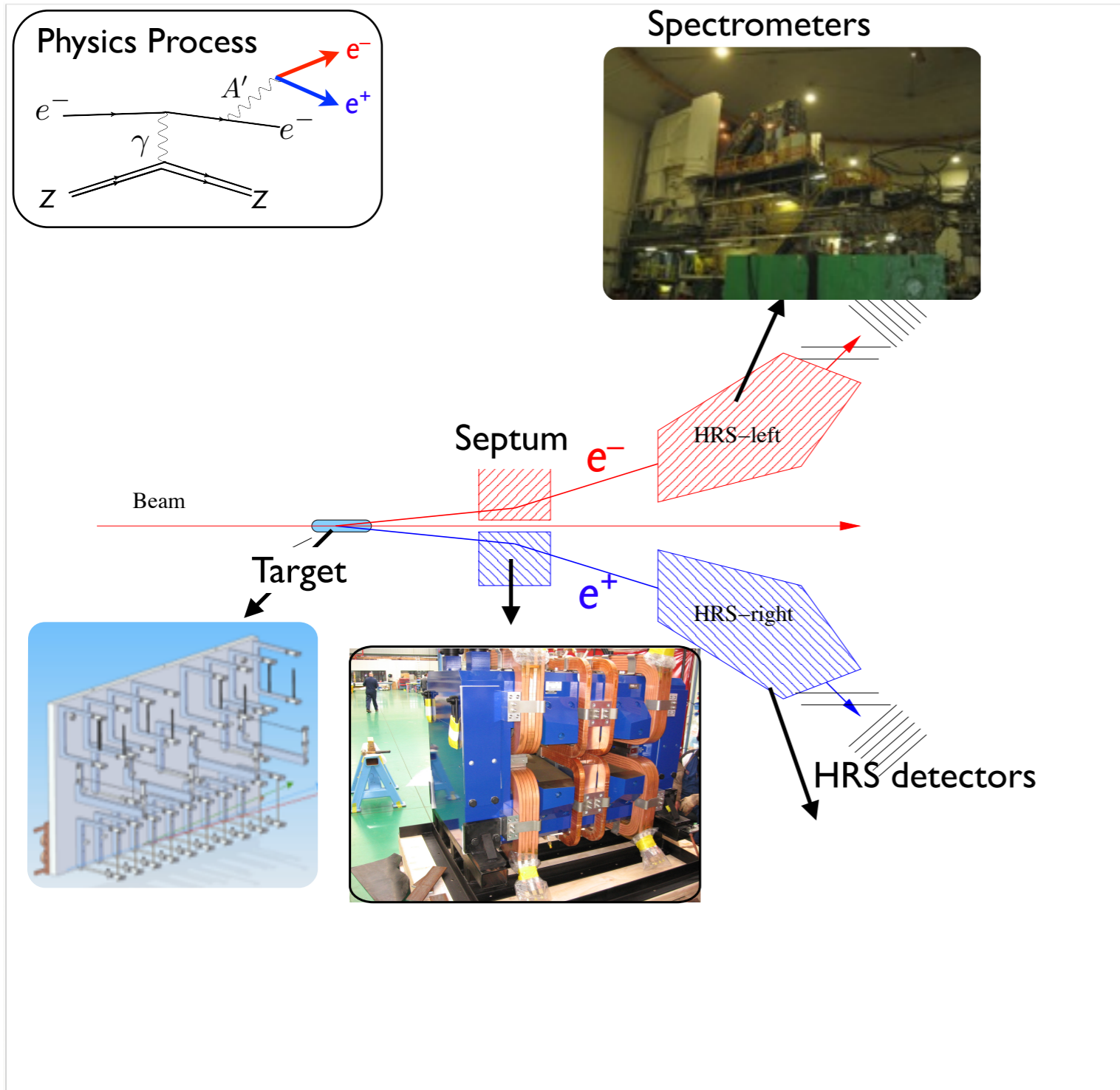




# APEX

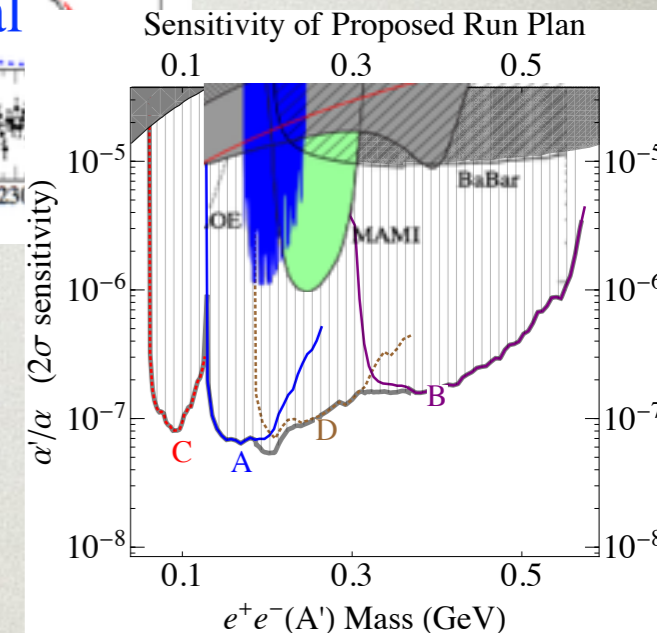
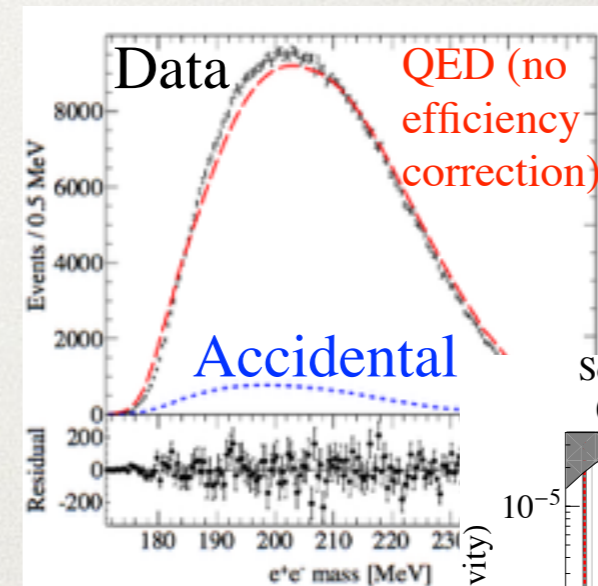
<http://hallaweb.jlab.org/experiment/APEX/>

Search for new gauge boson  $A'$  using Hall A high-resolution spectrometers (HRS)



## Status

**Test run (2010):** concept & technical demonstration; weekend run achieved world-record sensitivity

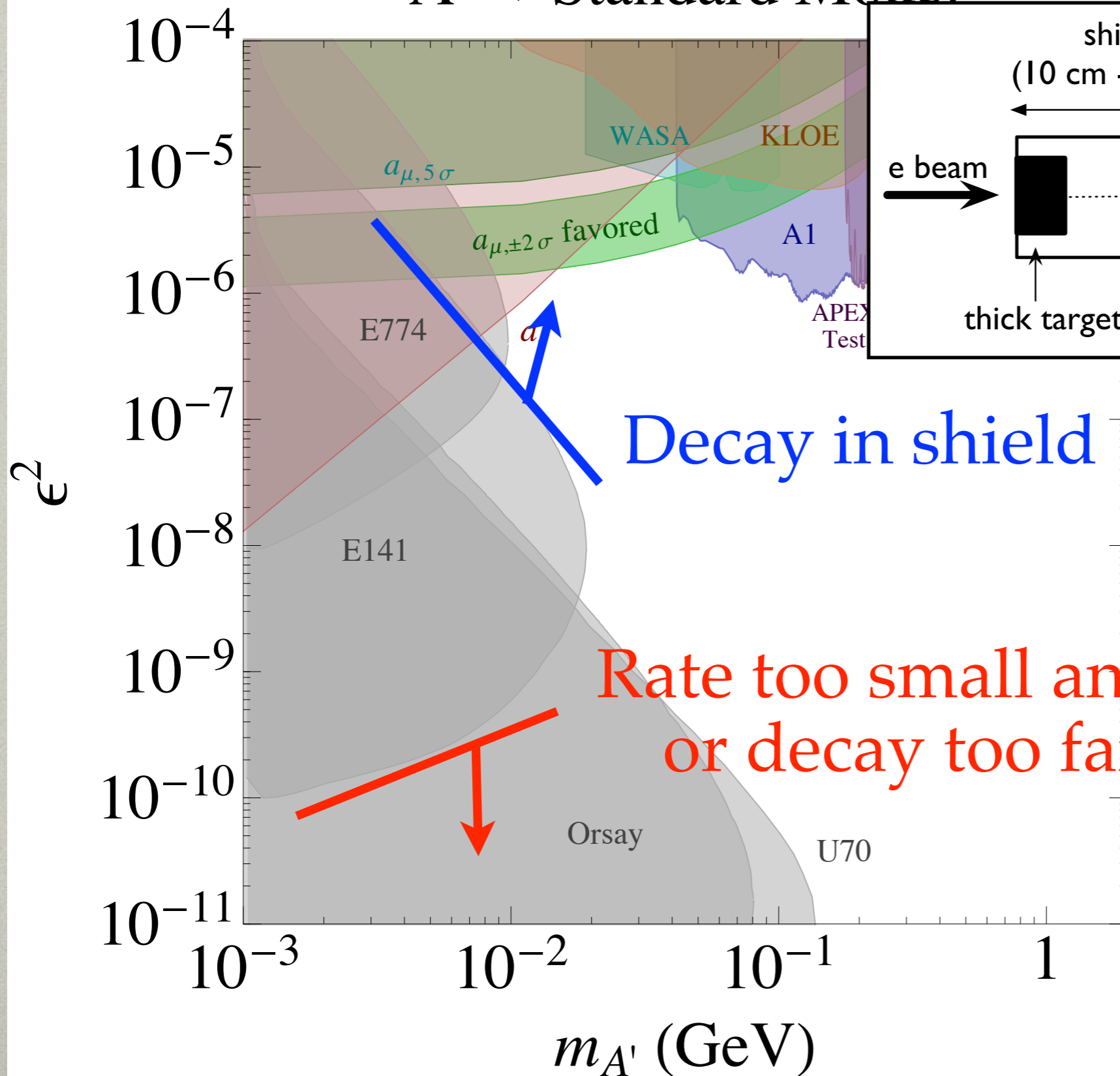


**Full run:** Prioritized by JLab, funded, projected to run 2017  
Optimized septa magnet constructed  
Smaller beam line items funded  
HRS detectors ready to go



# TURNING WEAKNESS INTO STRENGTH

$A' \rightarrow$  Standard Model

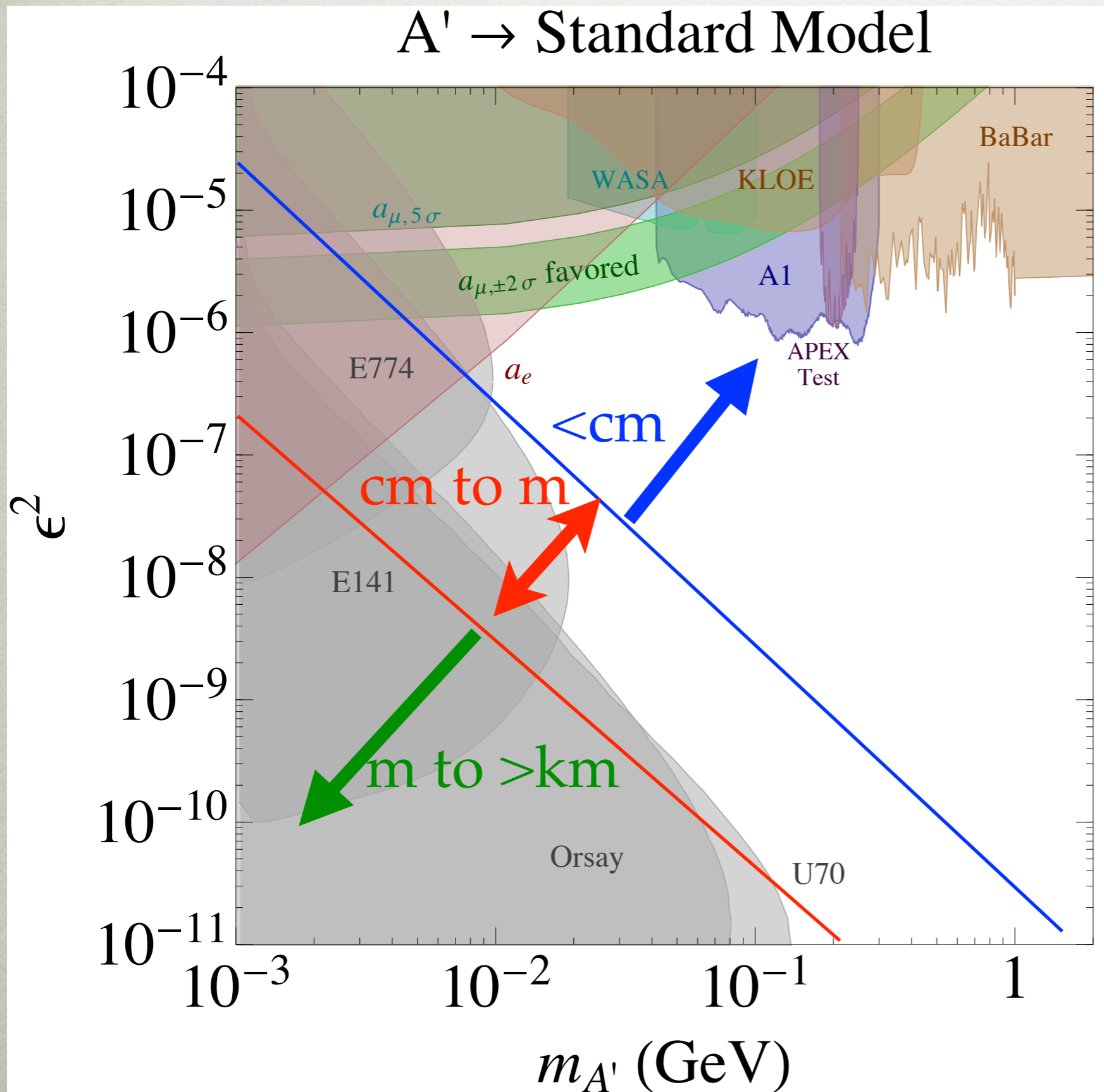


Lab-frame lifetime

$$\propto \frac{E}{m_{A'}^2 \epsilon^2}$$

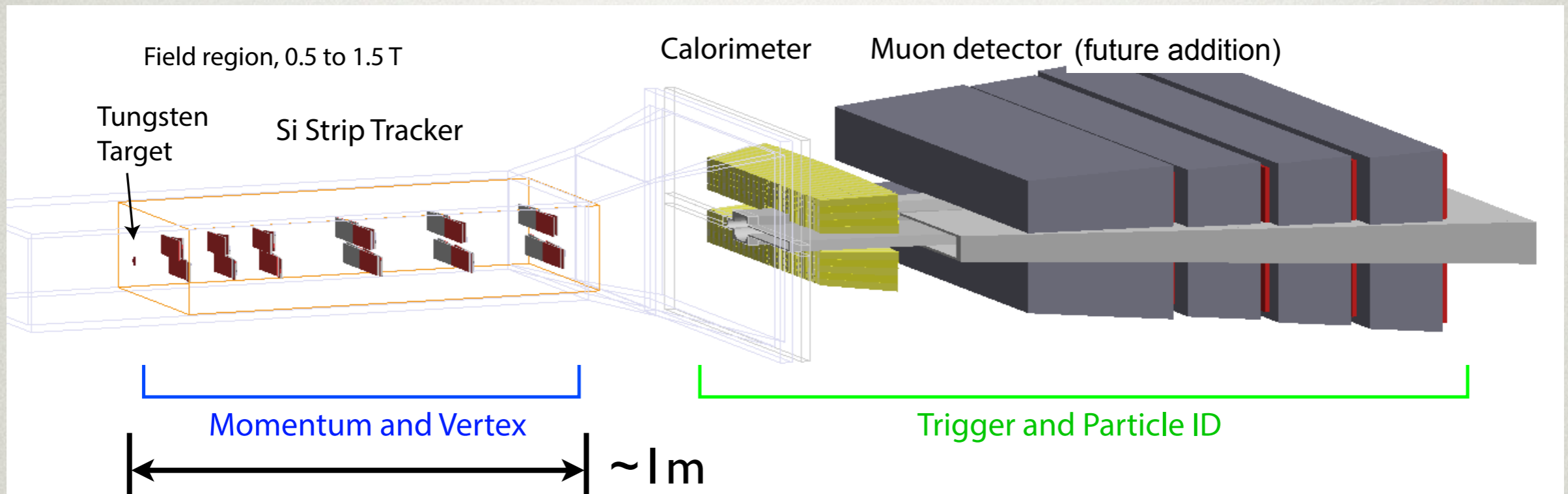


# TURNING WEAKNESS INTO STRENGTH

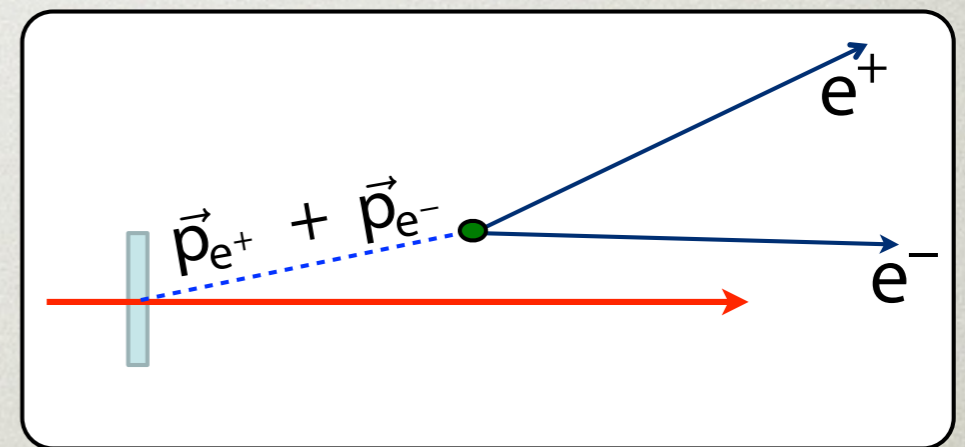




# HPS: RESONANCE + VERTEX SEARCHES



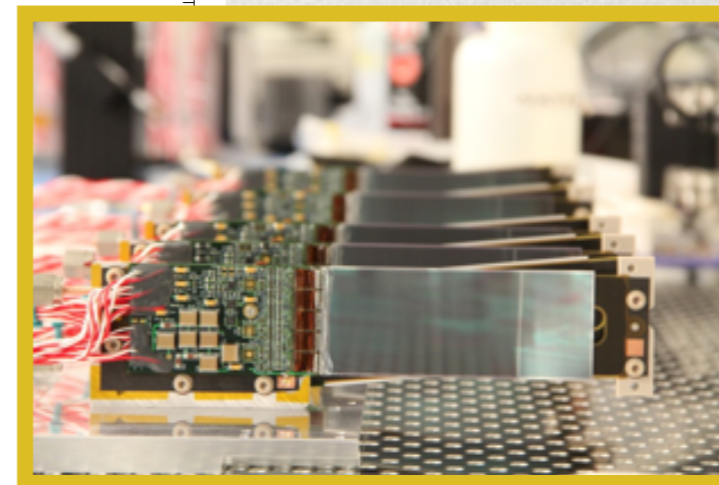
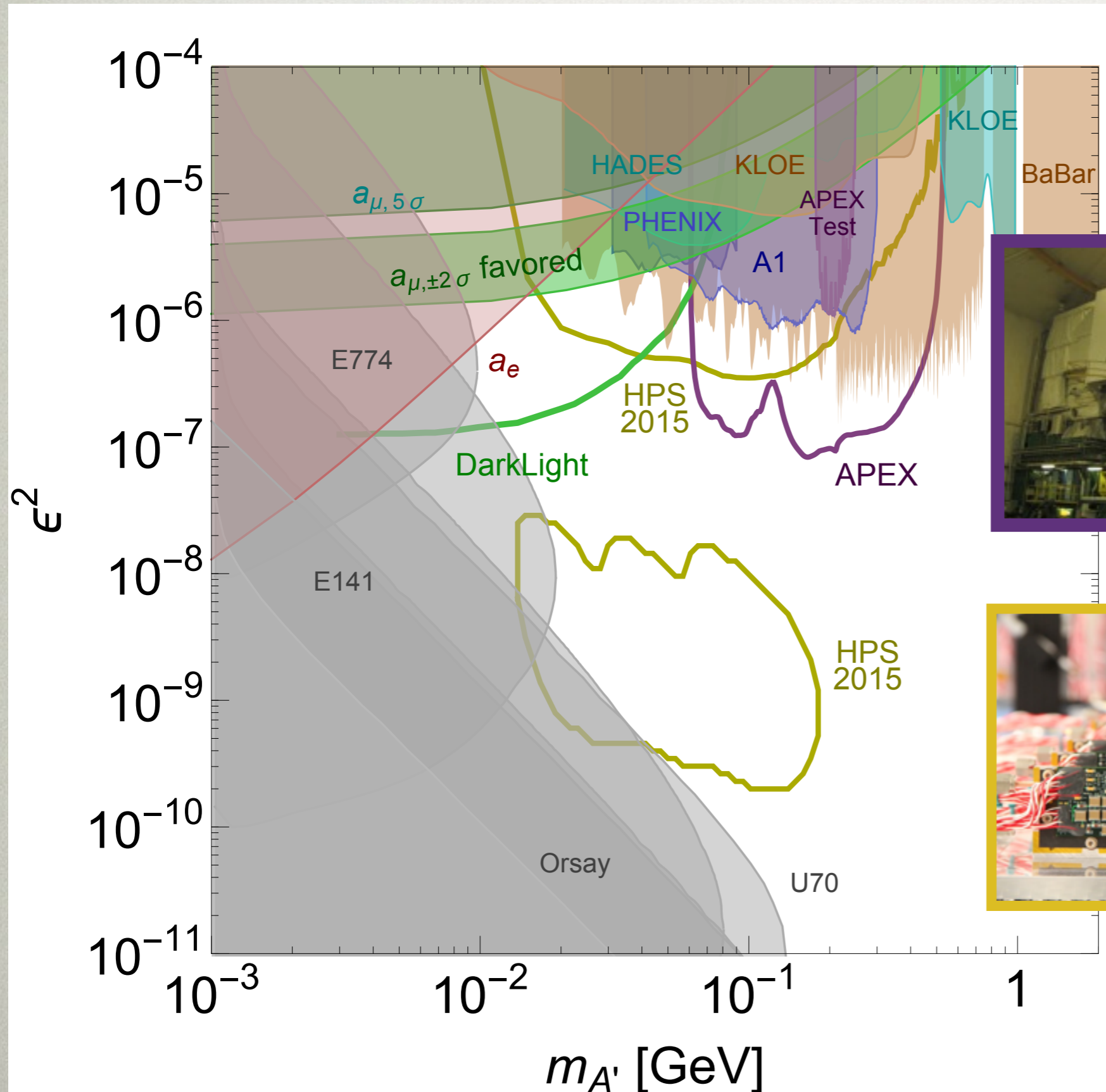
Allows sensitivity to very weak couplings with  $\sim$ cm decay vertex



**Engineering run this spring!**

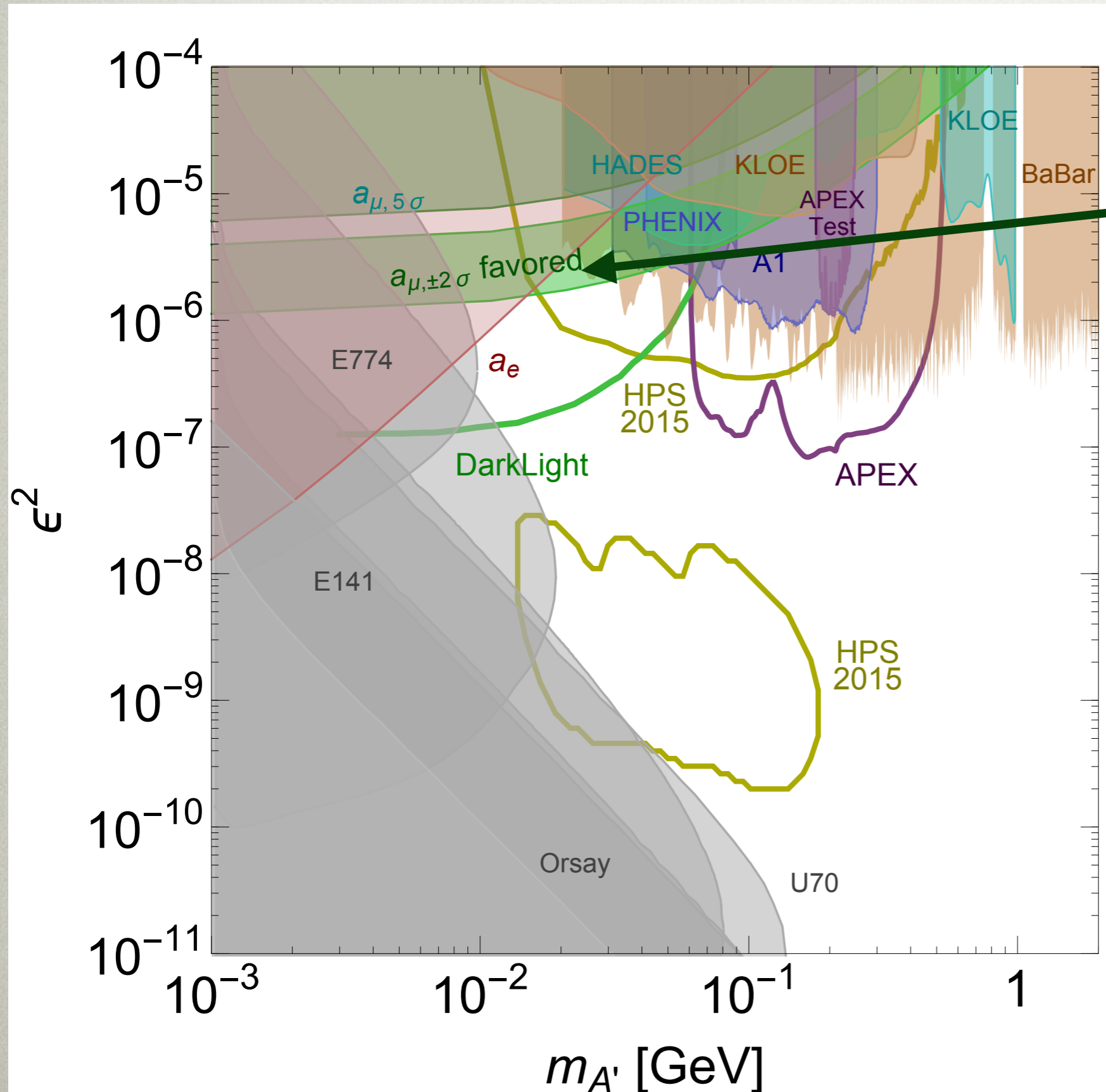


# PROJECTIONS FOR 2015-17





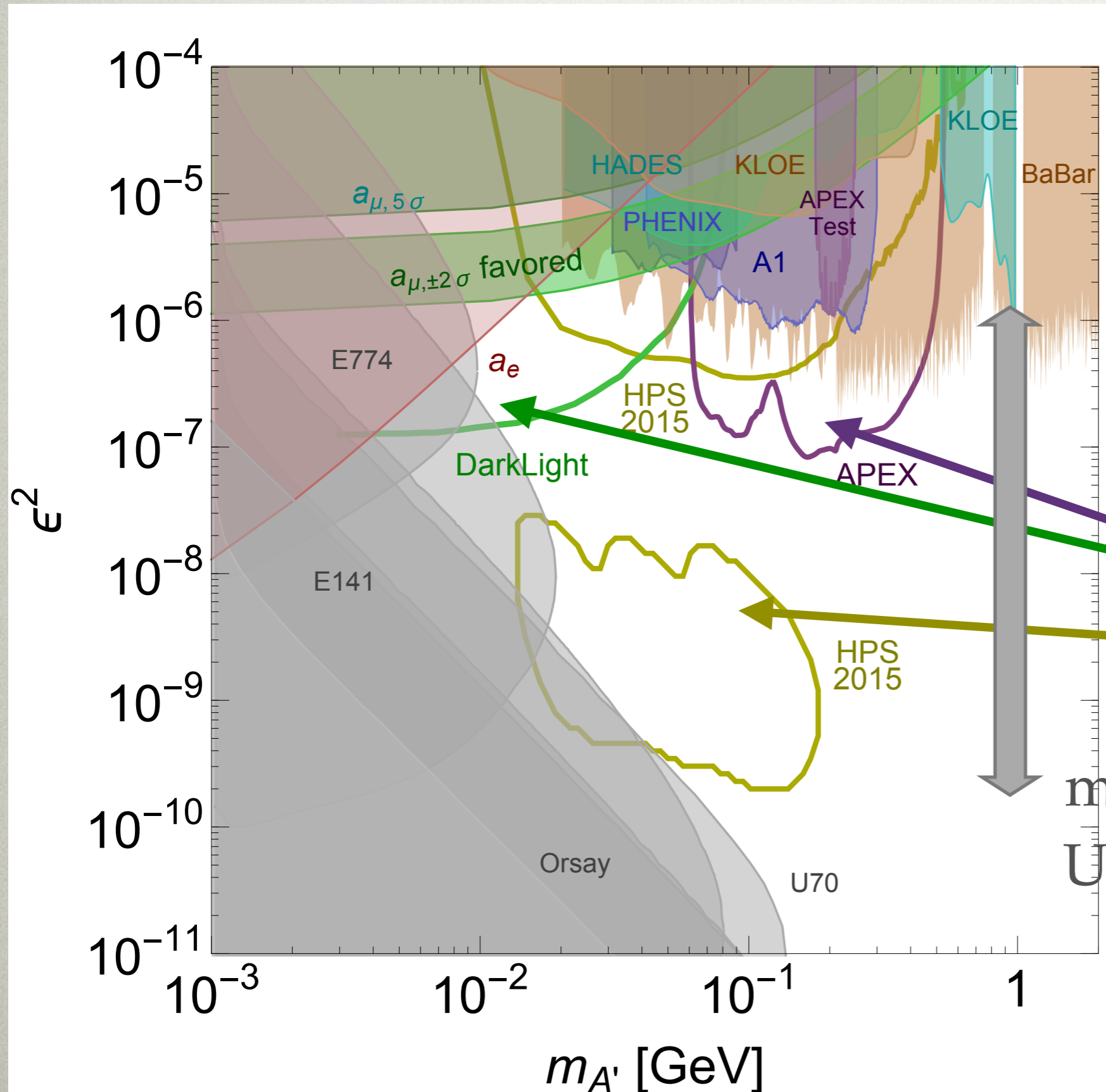
# PROJECTIONS FOR 2015-17



Tested this interpretation of muon magnetic moment anomaly!



# PROJECTIONS FOR 2015-17



Upcoming experiments explore this region from above:

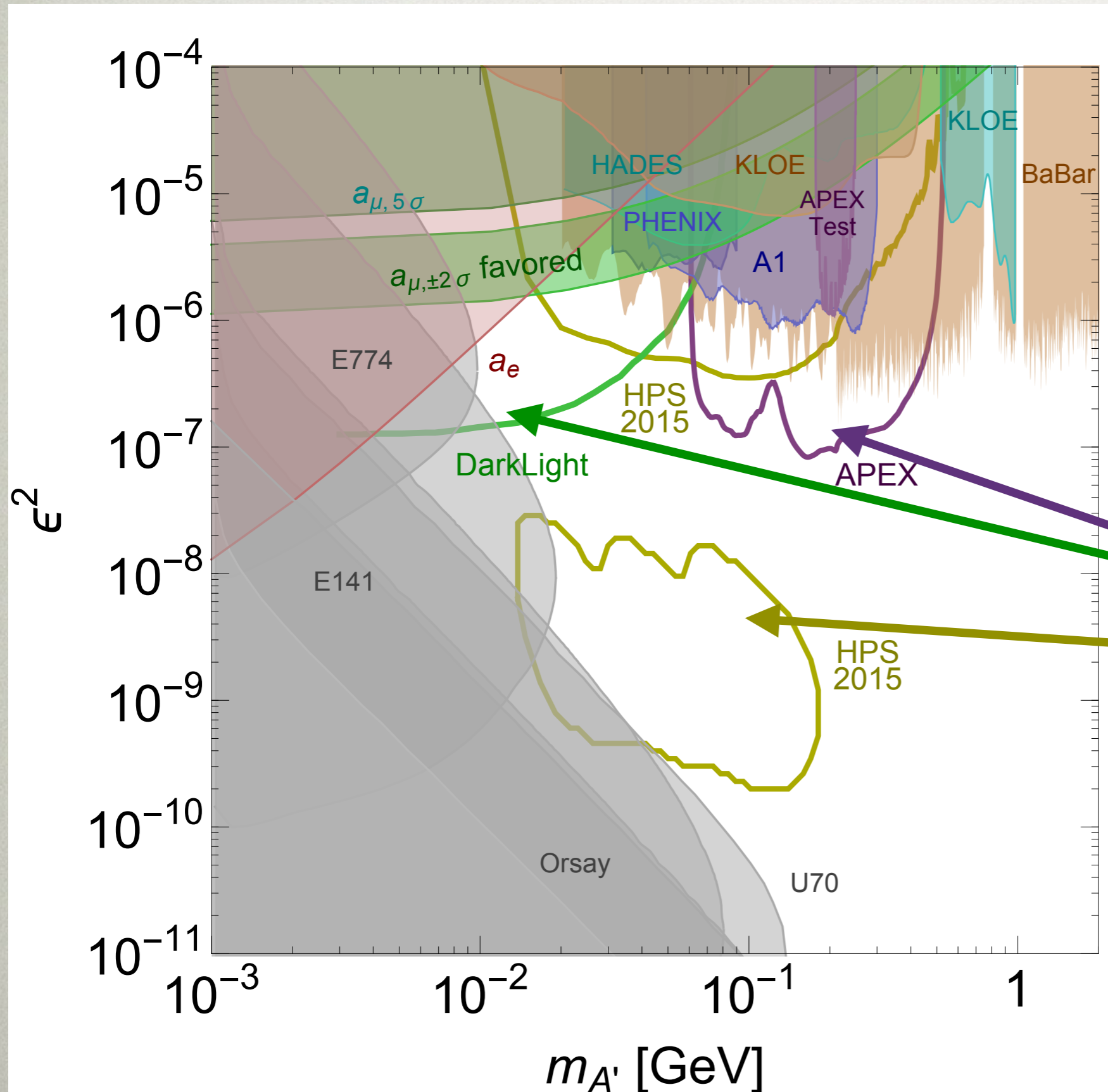
**APEX, DarkLight**

**and below: HPS**

mixing in Grand Unified Theories



# PROJECTIONS FOR 2015-17



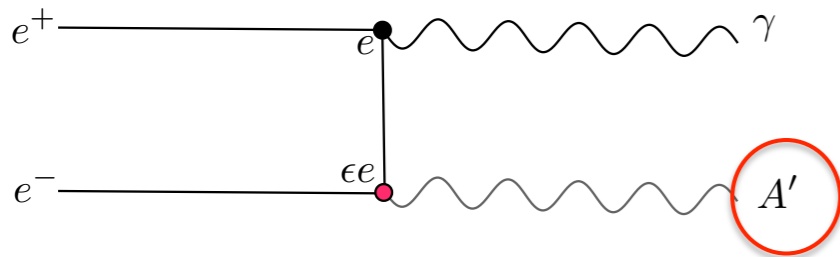
Also interesting  
mass range for  
dark matter  
interactions



# DARK PHOTONS AT CORNELL

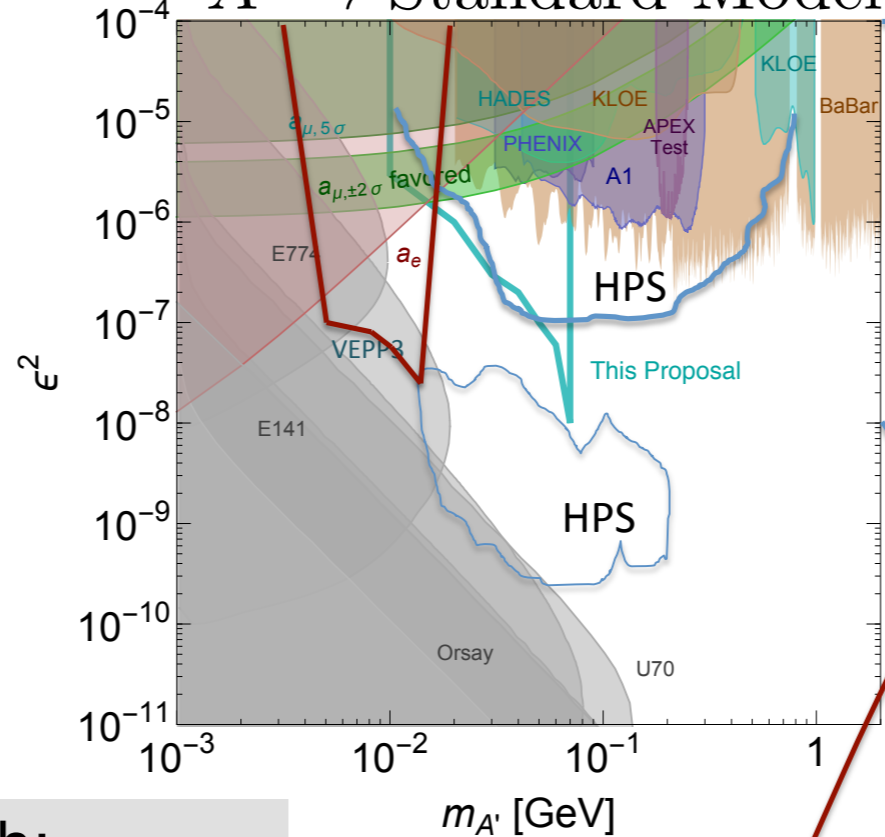
(figures from J. Alexander's talk to APEX collab. mtg, April 2015)

## Dark Model

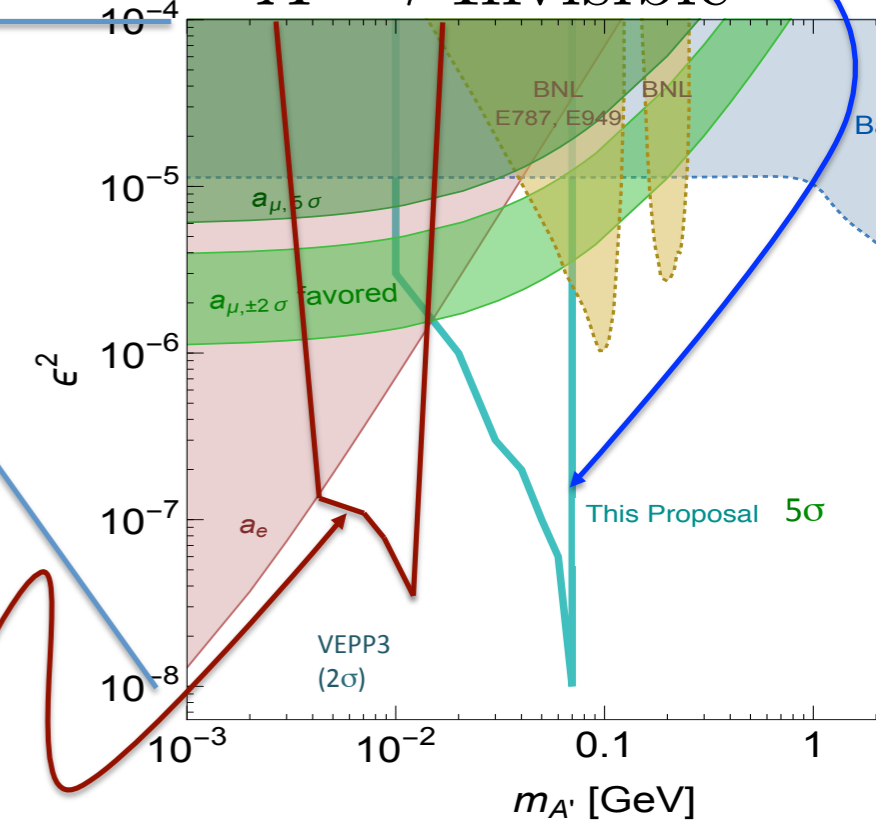


$E_{\text{beam}} = 5.3 \text{ GeV}$ ,  $I_{\text{beam}}^{\text{avg}} = 2.3 \text{ nA}$ ,  $\text{Lumi} = 1.0 \times 10^{34}$ ,  $T = 10^7 \text{ sec}$ , 5-sigma excl

## $A' \rightarrow \text{Standard Model}$

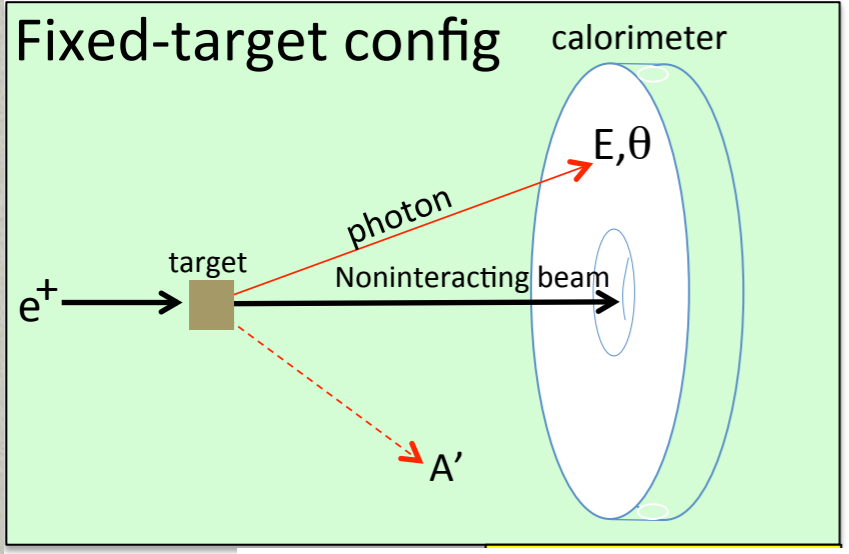


## $A' \rightarrow \text{Invisible}$



"Sister experiment" at Novosibirsk (proposed)

3



- **Missing mass** search:
  - Independent of  $A'$  decay mechanism
  - Bump hunt
  - Limited by  $\sqrt{s}$ 

$$M'_A < \frac{1}{2} \sqrt{s} \approx \sqrt{m_e E_{\text{beam}}/2}$$
    - 5 GeV beam:  $M'_A < 71 \text{ MeV}$
- Need positrons
- *This is the basic idea for the  $A'$  search at Cornell.*

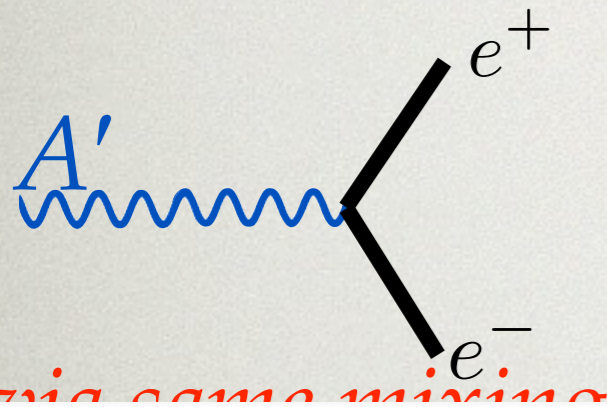
**Complementary in mass-coupling space and independent of  $A'$  decay mode**



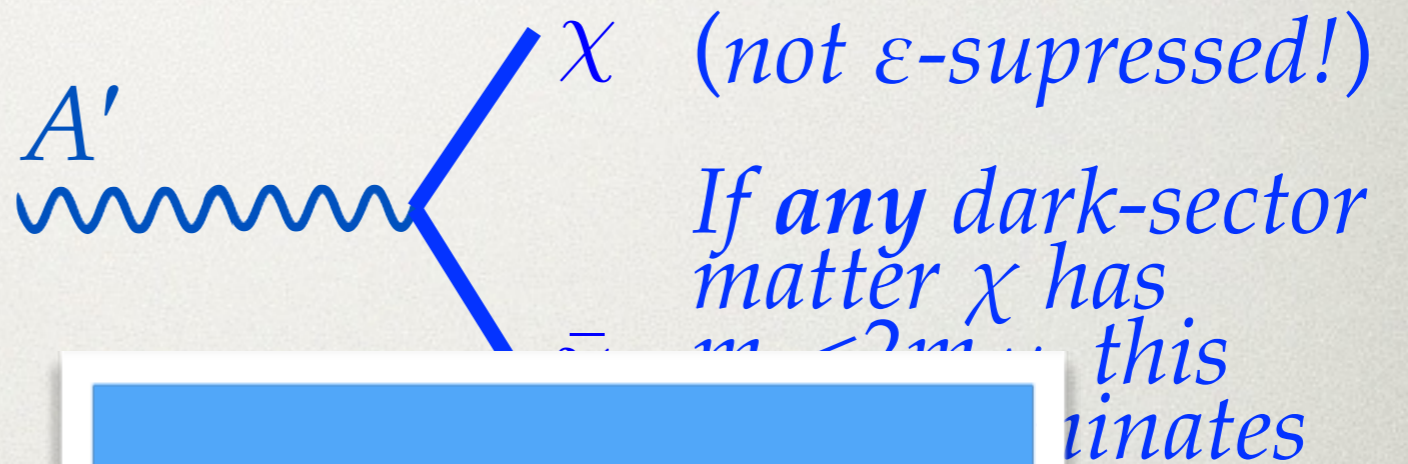
# A FIELD GUIDE TO DARK FORCES

## DECAY

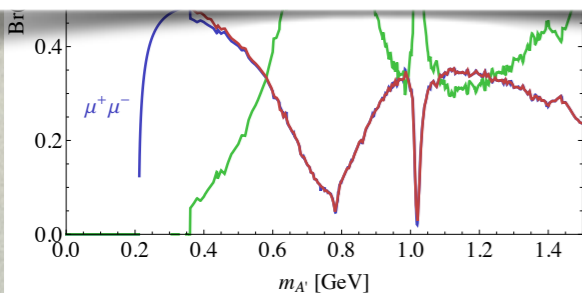
“Minimal” Decay:



“Generic” Decay:



Major advances in last 5 years!



Huge upcoming opportunity!

$T_W$

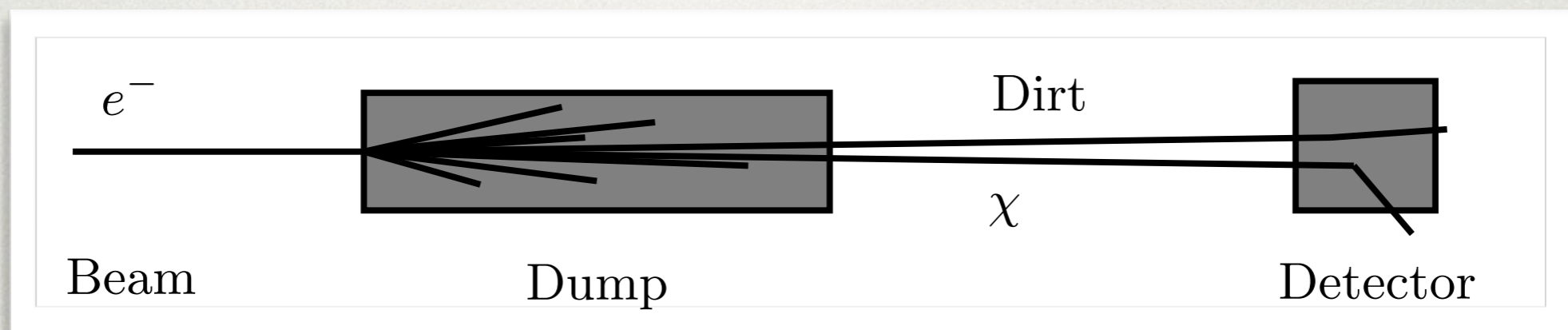
$A \rightarrow \gamma\gamma$  charged particles searches at BaBar and KLOE

To test “dark sector” idea, we need to search for both!

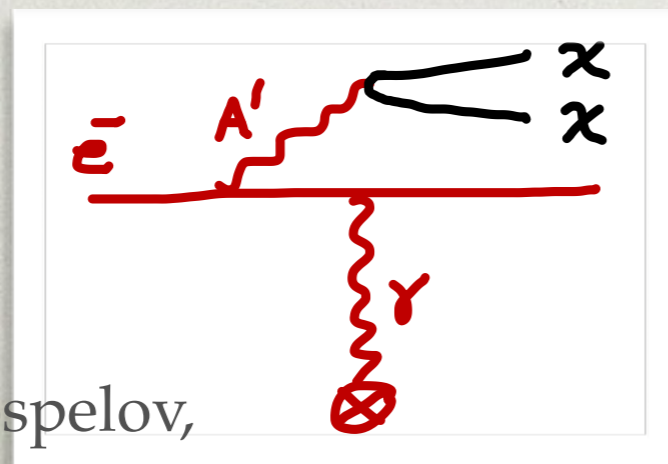


# DARK MATTER FROM THE DARK SECTOR

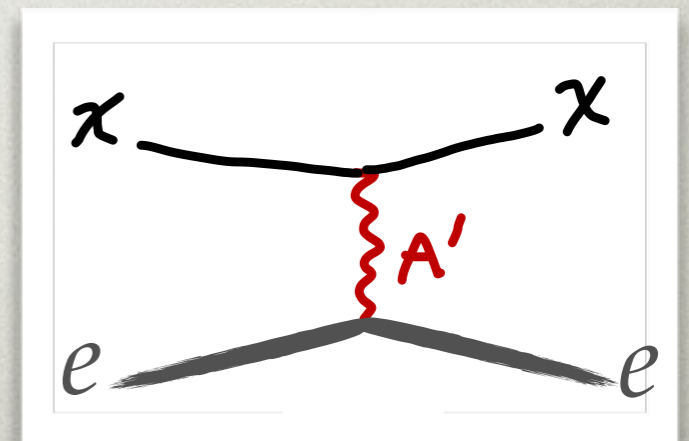
The same interaction that produces dark matter in the early Universe can also lead to production and detection in laboratory experiments!



Produce DM through the portal...



...detect its scattering downstream



Izaguirre, Krnjaic,  
Schuster & NT  
PRD.88.114015 and  
1403.6826

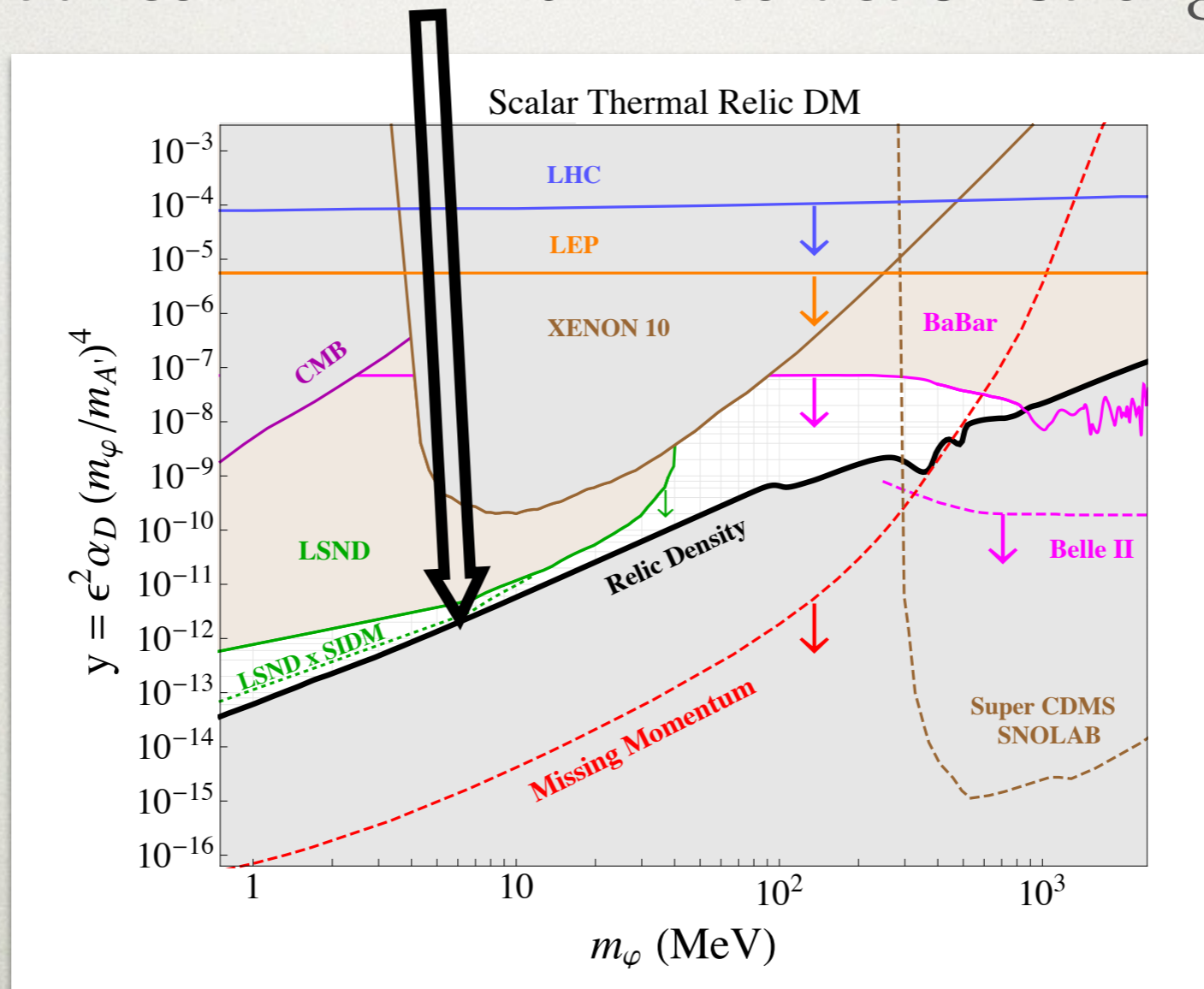
Proton beams: Batell,  
deNiverville, Ritz, Pospelov,  
McKeen, Dharmapalan, ...



# DARK MATTER FROM THE DARK SECTOR

The same interaction that produces dark matter in the early Universe can also lead to production and detection in laboratory experiments!

(thermal abundance  $\Rightarrow$  minimum interaction strength)

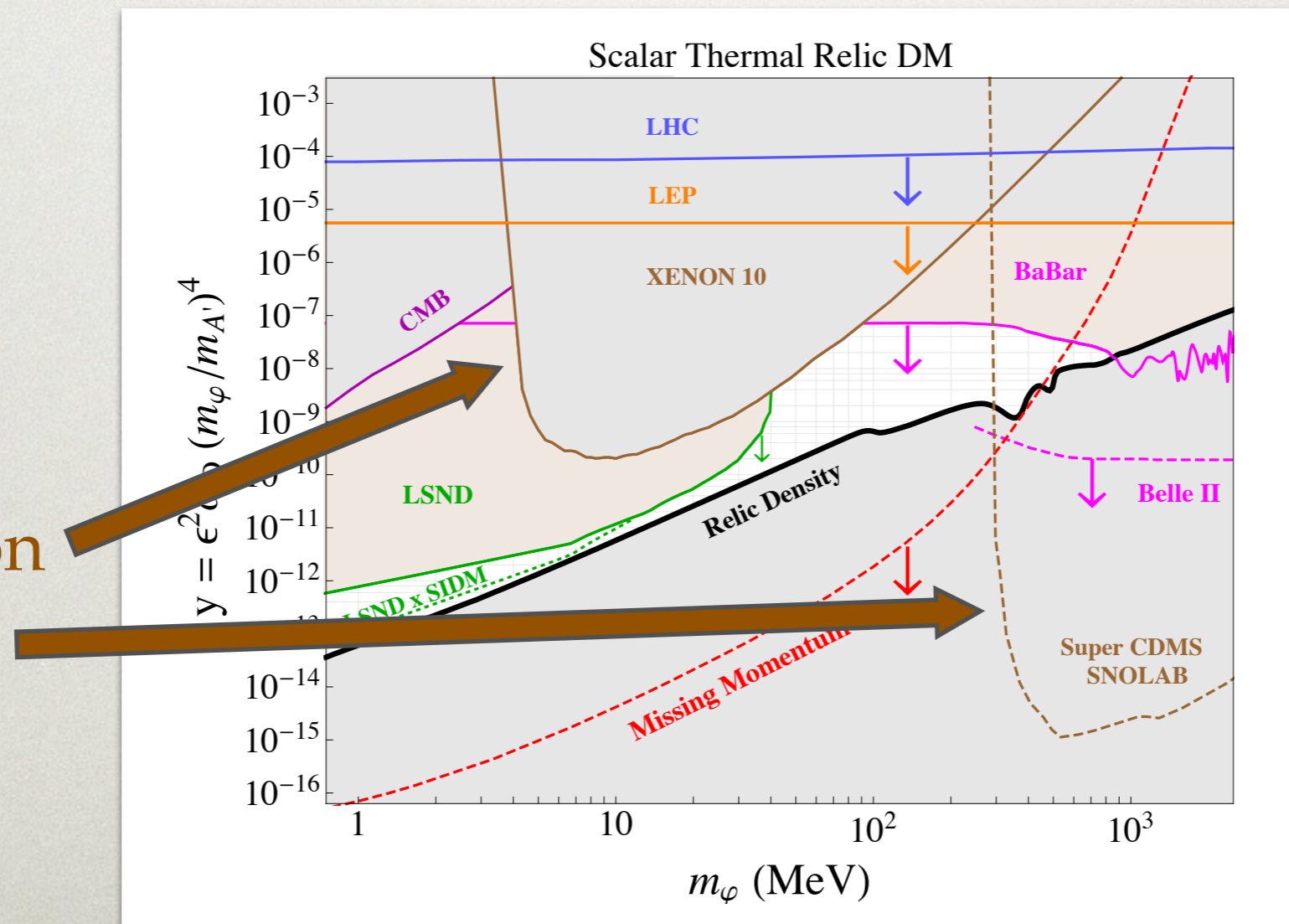




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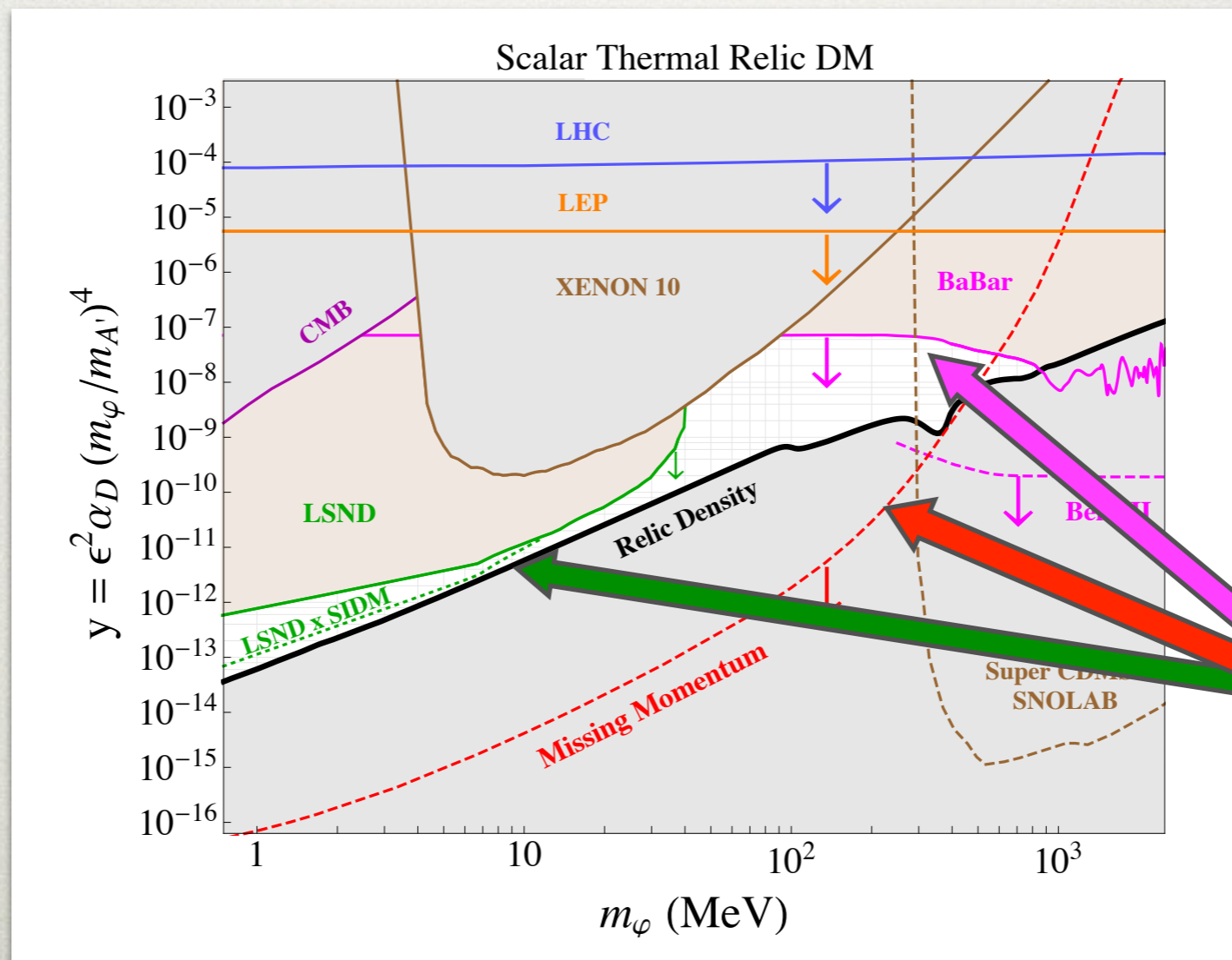
Hidden from direct detection in low-mass blind spot



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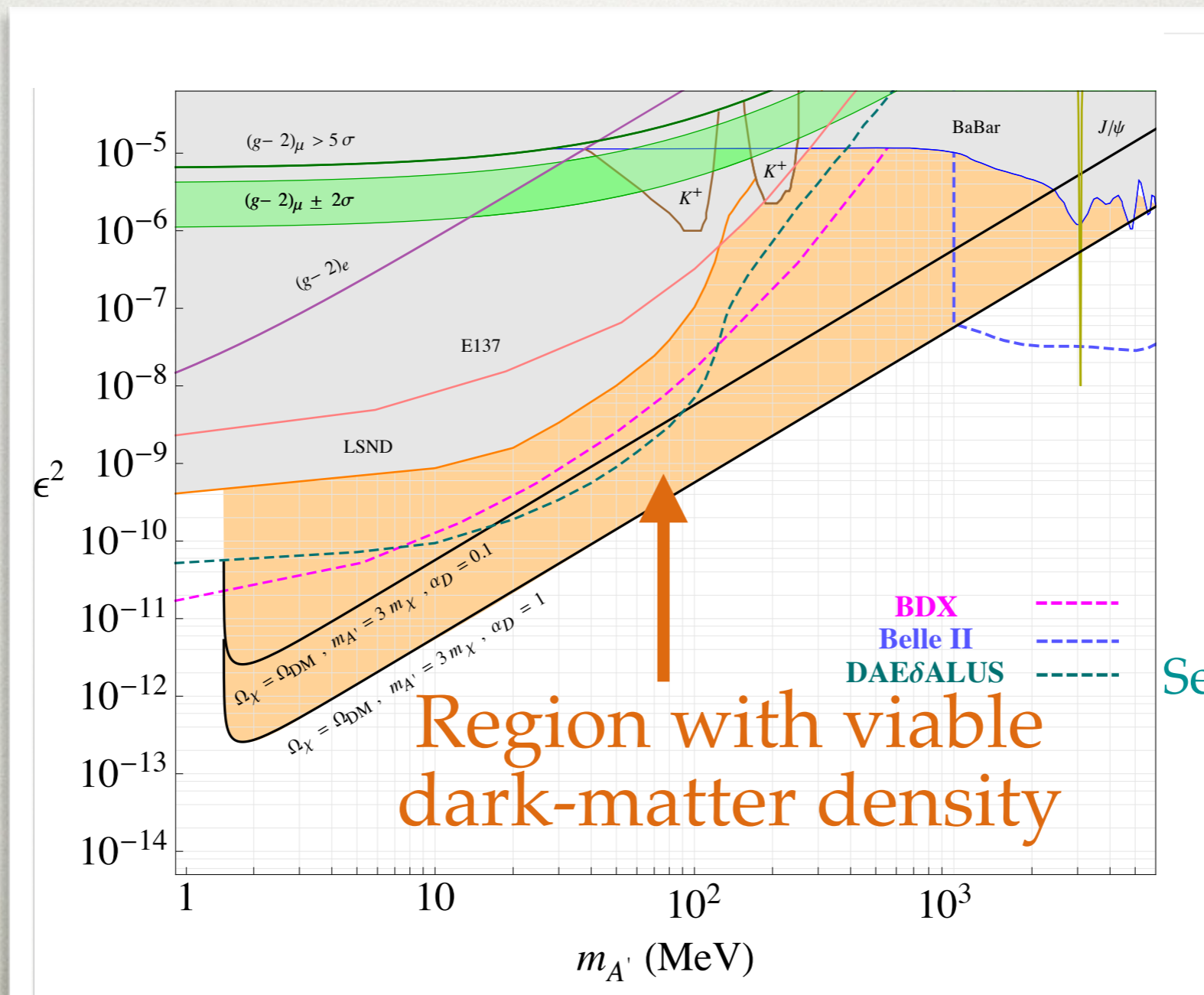
Hidden from direct detection in low-mass blind spot

More robustly tested by GeV-scale accelerator experiments



# DARK MATTER FROM THE DARK SECTOR

Directly in terms of the mixing  $\epsilon^2$ ...focus on low masses

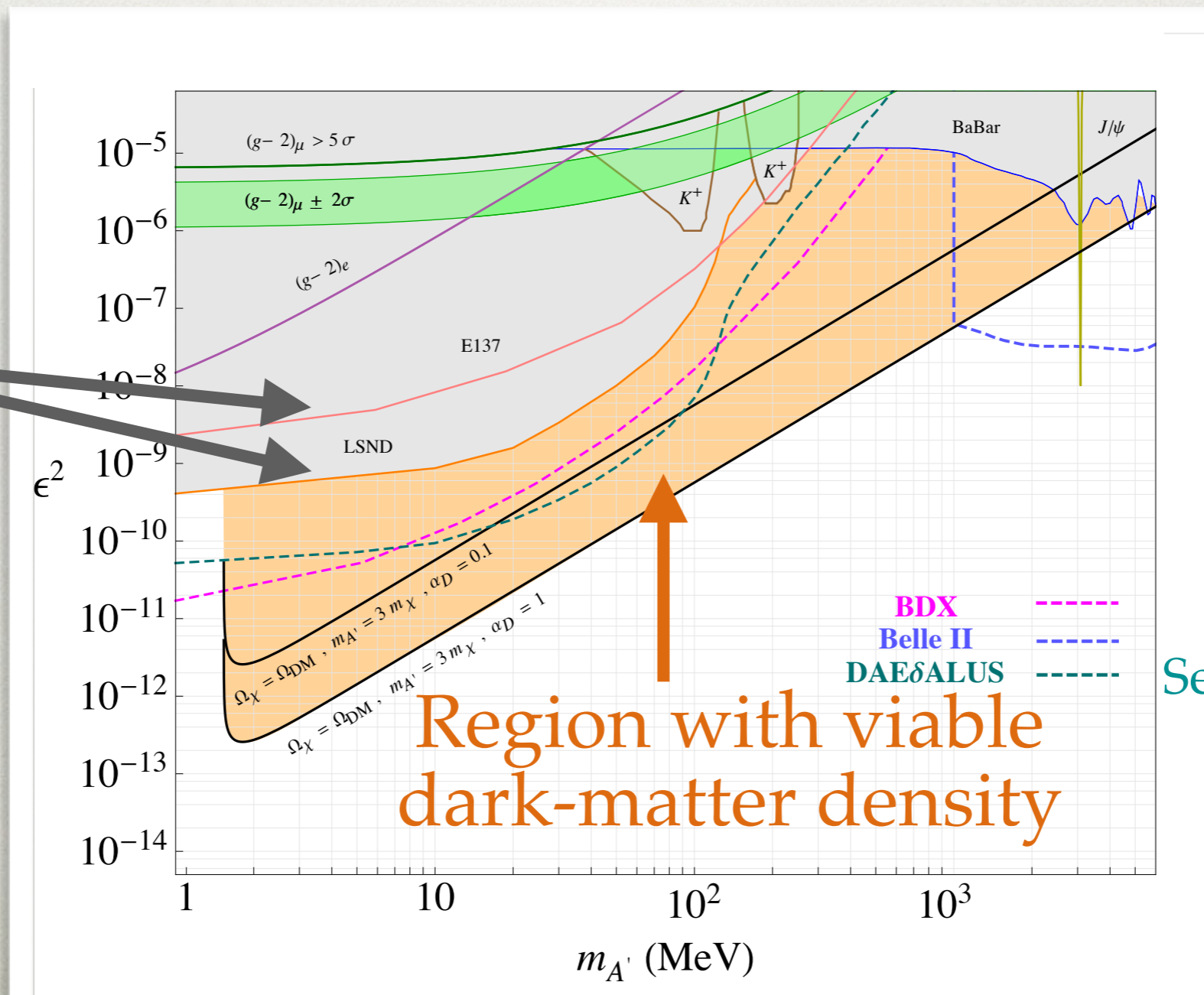




# DARK MATTER FROM THE DARK SECTOR

Directly in terms of the mixing  $\epsilon^2$ ...focus on low masses

Constraints from past  $e^-$  and  $p$  beam-dump experiments



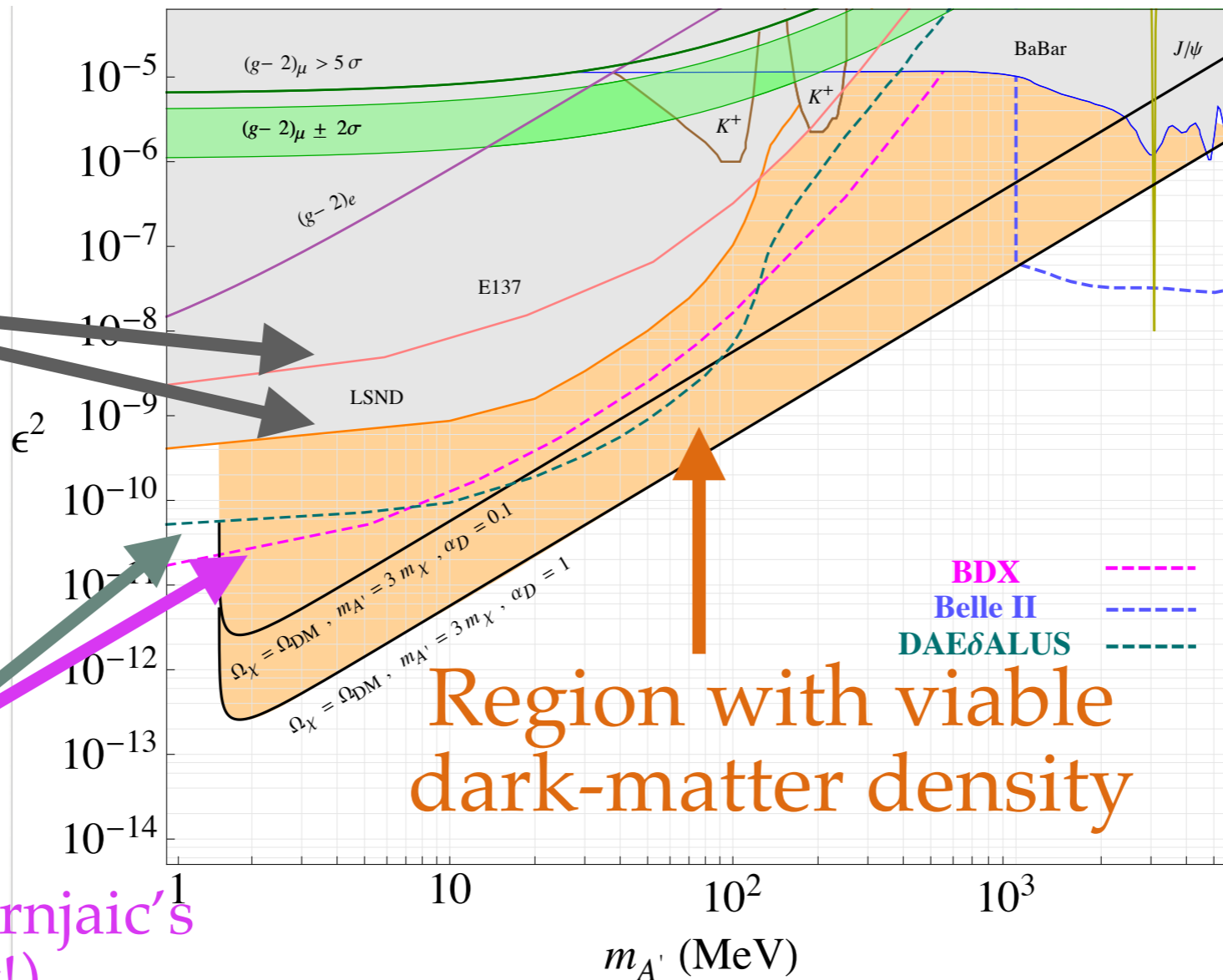
See Kahn, Krnjaic, Thaler, Tups 1411.1055



# DARK MATTER FROM THE DARK SECTOR

Directly in terms of the mixing  $\epsilon^2$ ...focus on low masses

Constraints from past  $e^-$  and  $p$  beam-dump experiments



Region with viable dark-matter density

Future prospects

(see Gordan Krnjaic's talk tomorrow!)

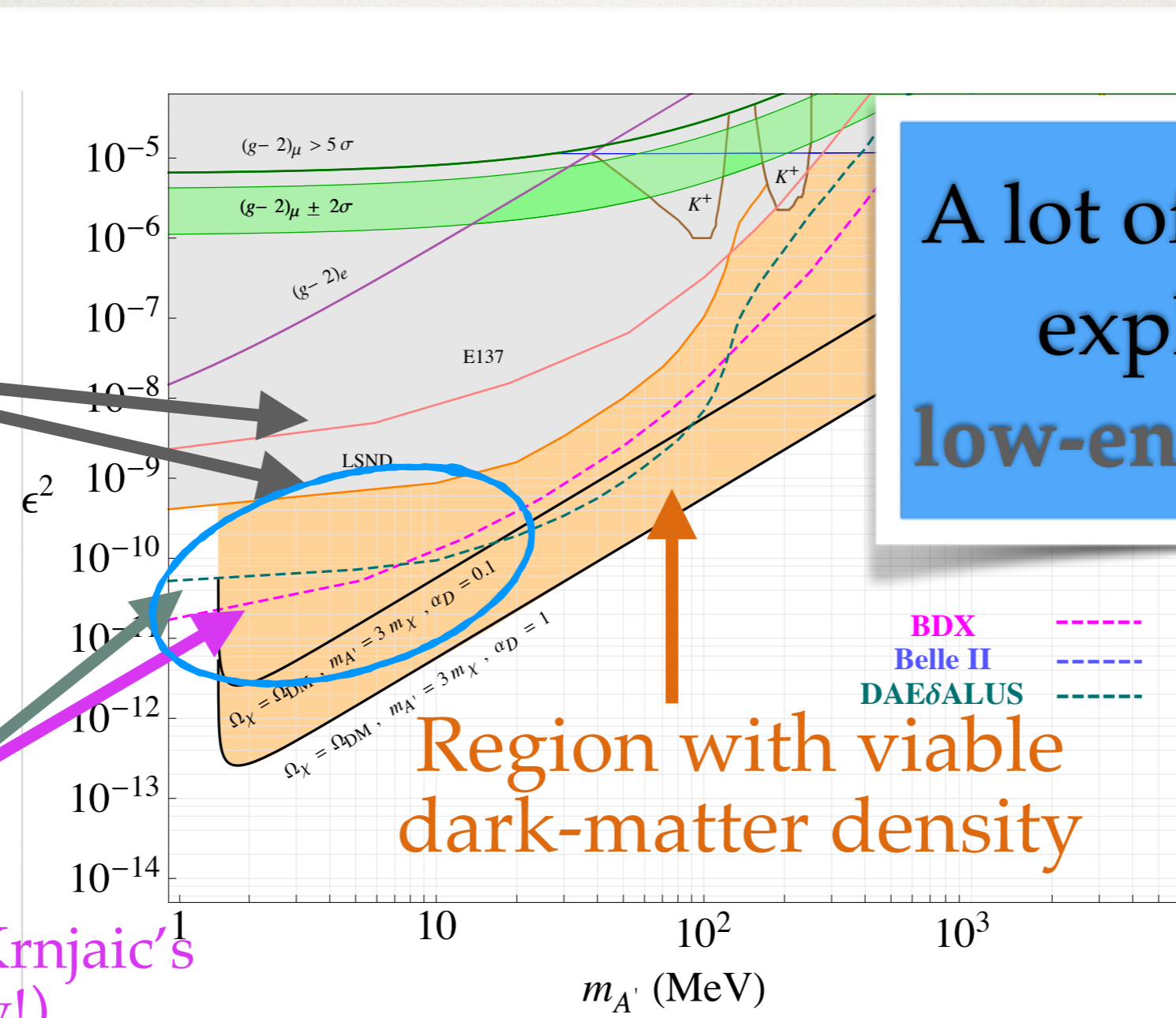
See Kahn, Krnjaic, Thaler, Tups 1411.1055



# DARK MATTER FROM THE DARK SECTOR

Directly in terms of the mixing  $\epsilon^2$ ...focus on low masses

Constraints from past  $e^-$  and  $p$  beam-dump experiments



A lot of territory to explore with low-energy beams!

Region with viable dark-matter density

Future prospects (see Gordan Krnjaic's talk tomorrow!)

See Kahn, Krnjaic, Thaler, Tups 1411.1055

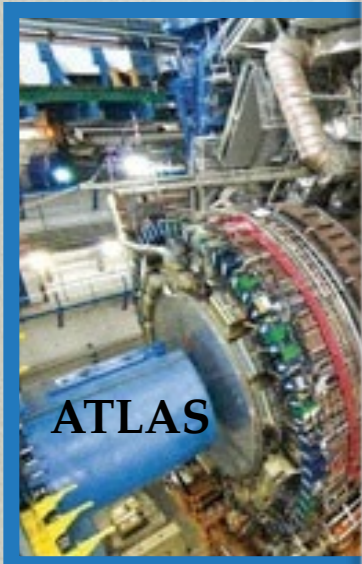


# THE FUTURE AHEAD

CMS



- Portals to explore physics neutral under Standard Model
  - Organize around **interaction** with ordinary matter, and visible vs. dark-sector decay
- Powerful sensitivity from current, planned, and ongoing experiments
- A lot of uncharted territory: opportunities abound for further exploration – and discovery – with intense electron beams!



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