

(light) **Dark Matter Overview**

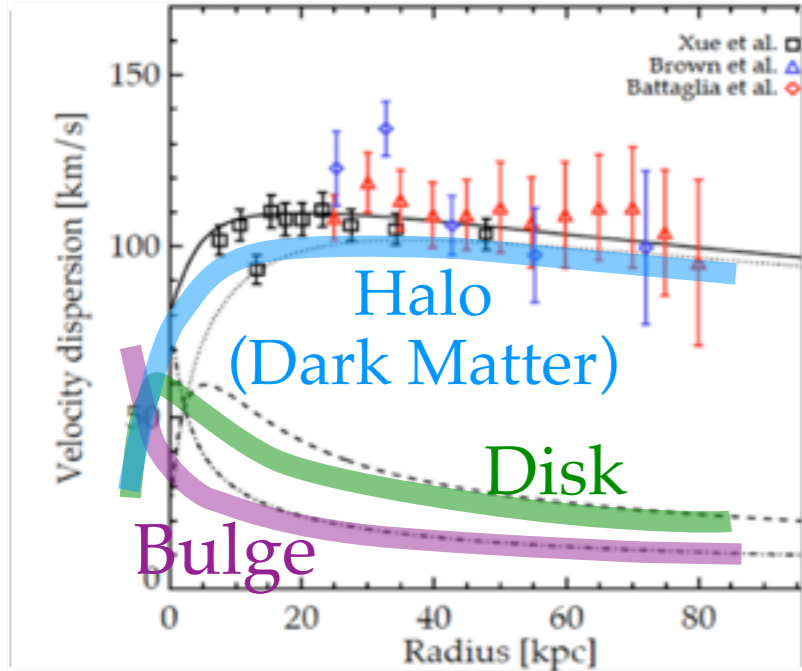
Gordan Krnjaic



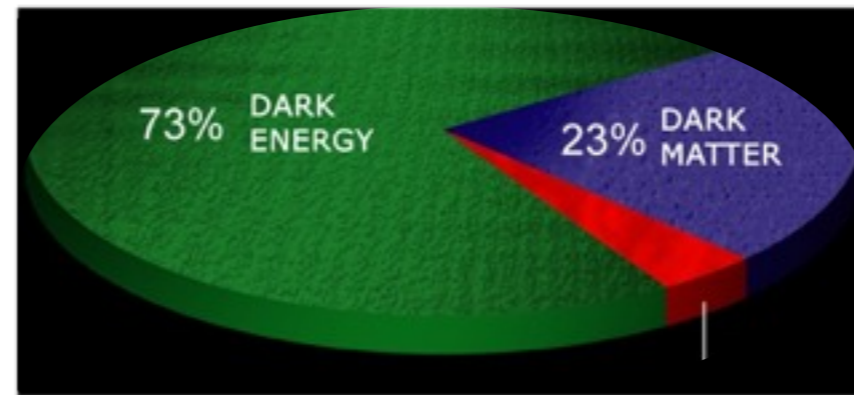
**Intense Electron Beam Workshop
Cornell University, June 18, 2015**

Impressive Indirect Evidence for DM

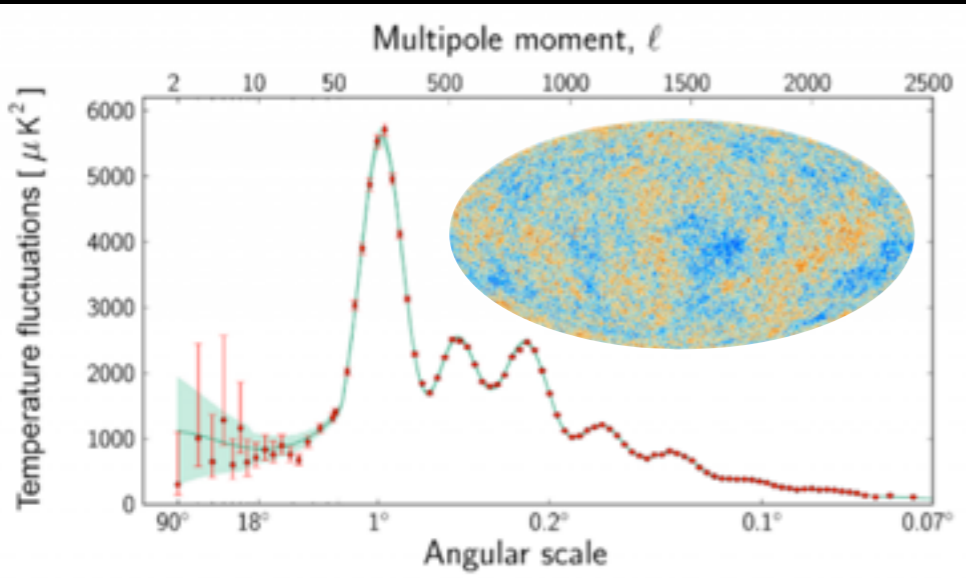
Rotation Curves



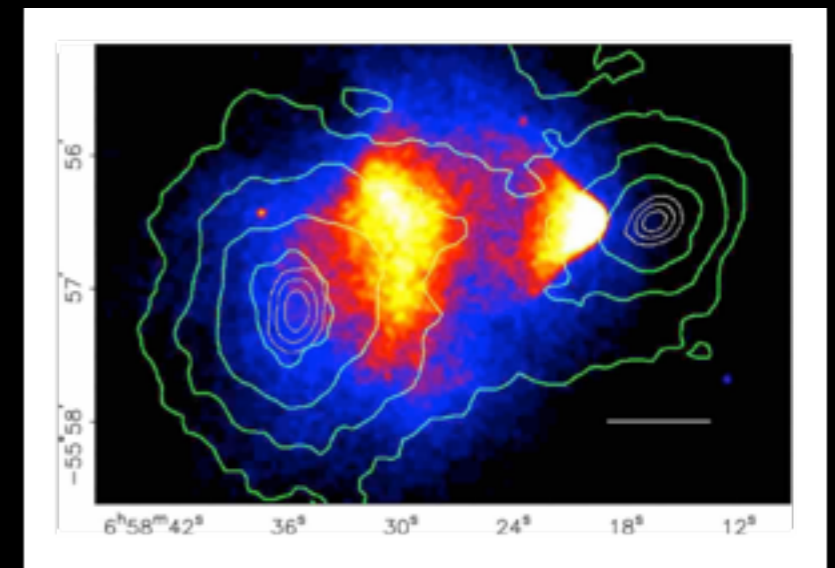
Gravitational lensing



CMB



Cluster collisions



but no direct observation (yet)!

Discovery challenge: no clear scale

DM viable over enormous mass range

$$10^{-33} \text{ eV} \longleftarrow m_{\text{DM}} \longrightarrow 10^{19} \text{ GeV}+$$

Hubble sized
axion-like particle

Black hole/MACHO

Many scenarios are undiscoverable

Fortunately, “thermal” history narrows scope

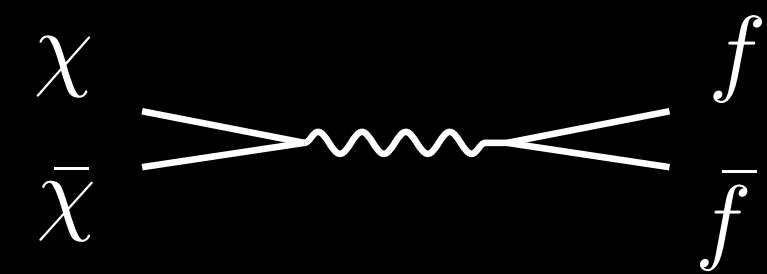
Applies to most* interesting / discoverable models
& lots of room for progress

Thermal Origin is Predictive

Feature # 1: huge early universe density!

$$n_{\text{DM}}(T) = \int \frac{d^3p}{(2\pi)^3} \frac{g_i}{e^{E/T} \pm 1} \sim T^3$$

Requires minimum annihilation rate

$$\sigma v \geq \sigma v_{\text{relic}} \sim 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$$


The diagram shows a particle χ and its antiparticle $\bar{\chi}$ on the left, meeting at a vertex. A wavy line represents the annihilation process. On the right, two particles f and \bar{f} emerge from another vertex.

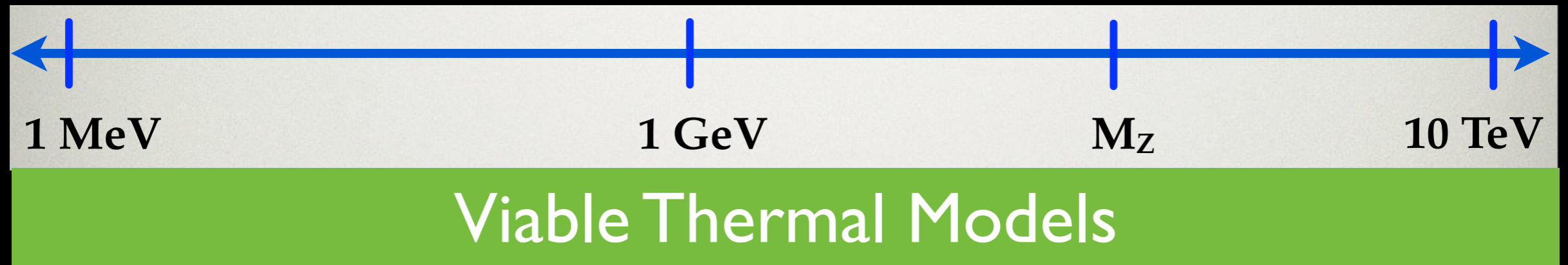
= symmetric DM

> asymmetric or subdominant DM

Important target for discovery or falsifiability

Thermal Origin is Predictive

Feature # 2: most masses can't be thermal



<10 keV DM too hot
spoil structure formation

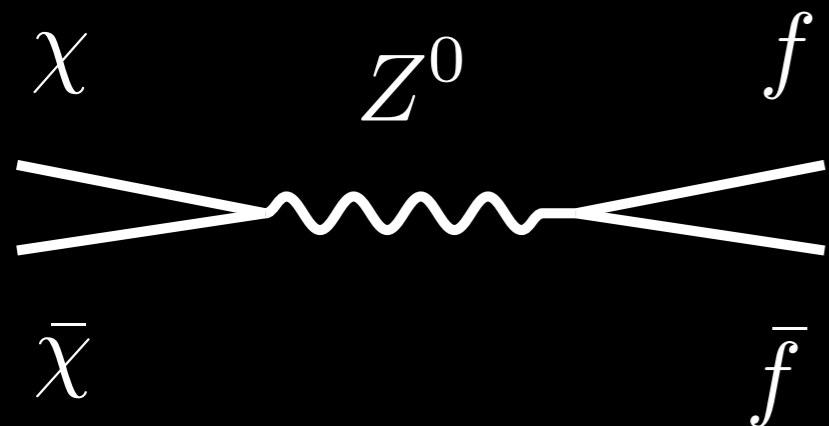
> 10 TeV DM
overproduced
and/or nonperturbative

Equilibrium reduces viable mass & coupling range

Annihilation “mediators”

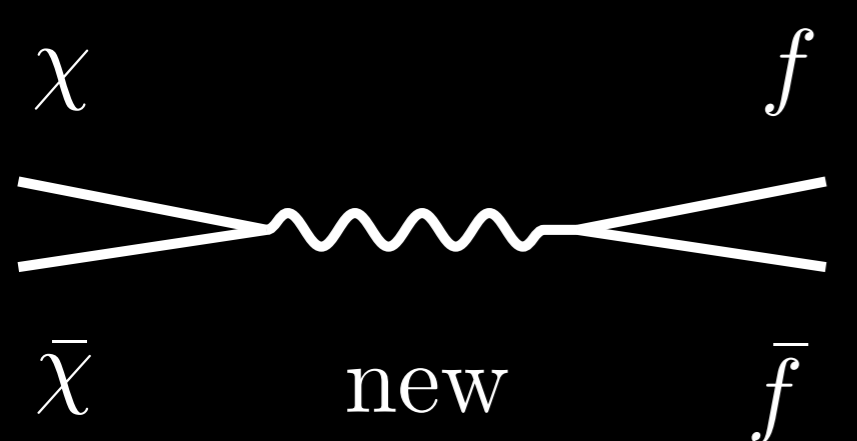
> GeV DM : Mediator can carry SM charge

“wimp miracle”


$$\sigma v \propto \frac{m_\chi^2}{m_Z^4} \implies \Omega_\chi \sim \Omega_{DM}$$

< GeV DM : W/Z/H too heavy!

“wimpless miracle”


$$\sigma v \propto \frac{m_\chi^2}{m_{\text{new}}^4} \implies \Omega_\chi \sim \Omega_{DM}$$

\implies Need light new mediator

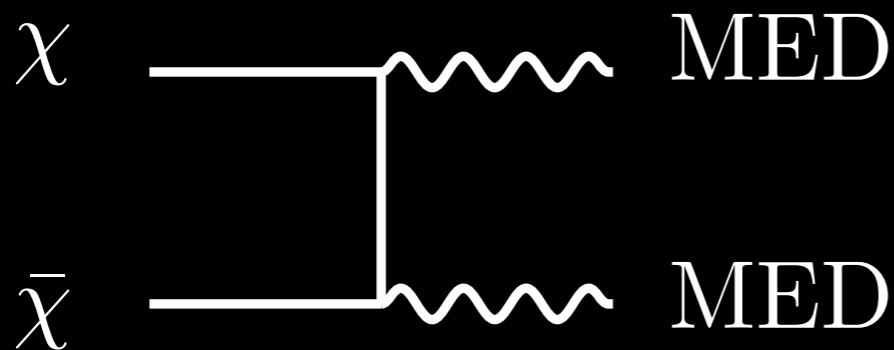
What kind of new “mediator”?

Direct annihilation = invisibly decaying mediator

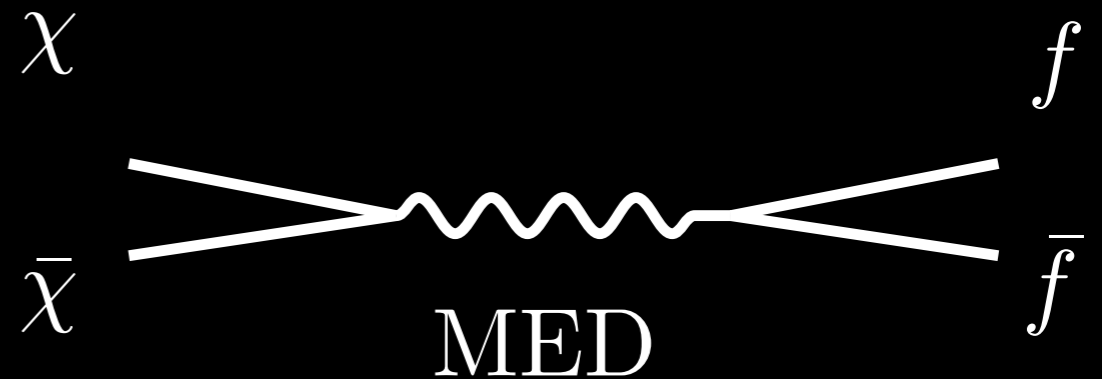
$$m_{\text{MED}} > 2m_{\text{DM}}$$

offers clear & testable target

Visible mediator decay
t-channel DM annihilation



Invisible mediator decay
s-channel DM annihilation



insensitive to SM coupling

sensitive to DM x SM coupling

What kind of new “mediator”?

Direct annihilation = invisibly decaying mediator

$$m_{\text{MED}} > 2m_{\text{DM}}$$

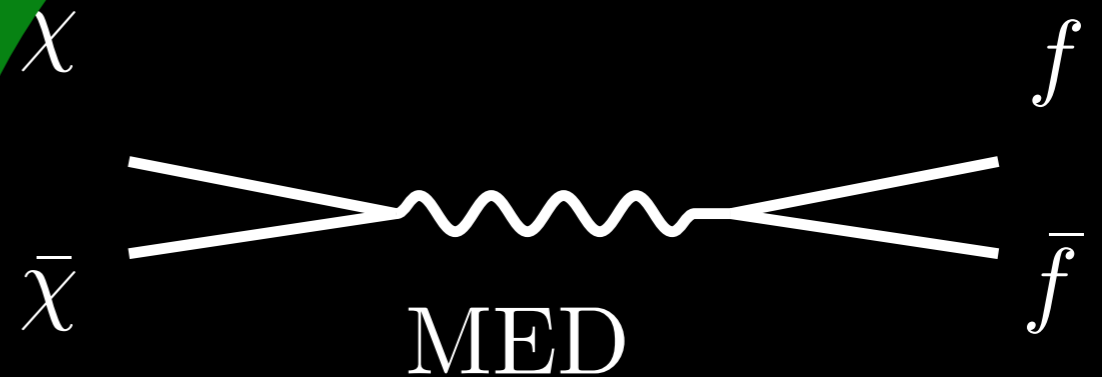
offers clear & testable target

Visible mediator decay
t-channel DM annihilation



insensitive to SM coupling

Invisibly decaying mediator decay
t-channel DM annihilation



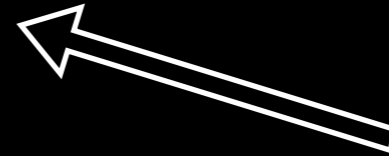
sensitive to DM x SM coupling

What kind of new “mediator”?

Light mediator must be SM neutral

Higgs Portal

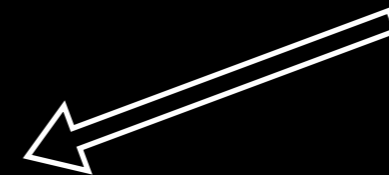
$$(H^\dagger H) |\phi^\dagger \phi|$$



couplings scale with mass

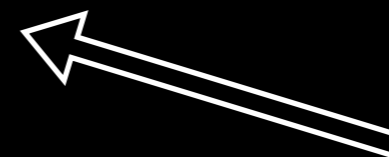
“Axion” Portal

$$\frac{m_f}{f_a} a \bar{f} \gamma^5 f$$



Vector Portal

$$\epsilon F_{\mu\nu} F'_{\mu\nu}$$



couplings scale with charge

**Invisibly decaying
A' mediator**

$$\mathbf{A}' \text{ to SM: } \epsilon e$$

$$\mathbf{A}' \text{ to DM: } g_D$$

Some caveats (see Natalia Toro's talk)

What kind of new “mediator”?

Higgs Portal

$$\langle H^\dagger H \rangle |\phi^\dagger \phi|$$

← Thermal Ruled out by rare meson decays

“Axion” Portal

$$\frac{m_f}{f_a} a \bar{f} \gamma_5 f$$

← [Izaguirre, GK, Schuster, Toro]
(to appear)

Vector Portal

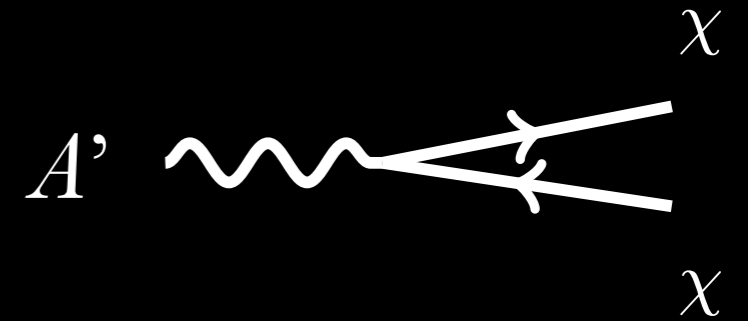
$$\epsilon F_{\mu\nu} F'_{\mu\nu}$$

← Many viable sub-GeV scenarios

What kind of DM?

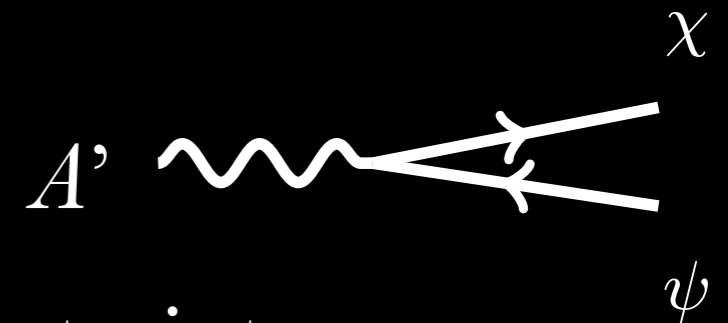
Classify by Coupling/Spin/Abundance

Elastic (no mass splitting)



Constrained by CMB & Direct Detection

Inelastic (mass splitting)



Easily evades CMB & Direct Detection constraints

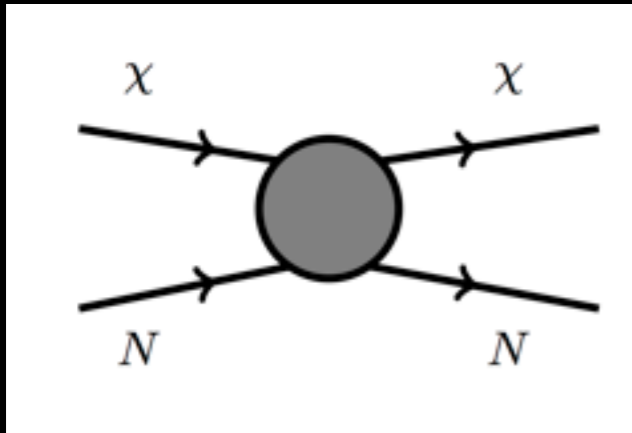
$$\Delta \equiv m_\psi - m_\chi$$

Either can be (fermion/scalar) x (symmetric/asymmetric)

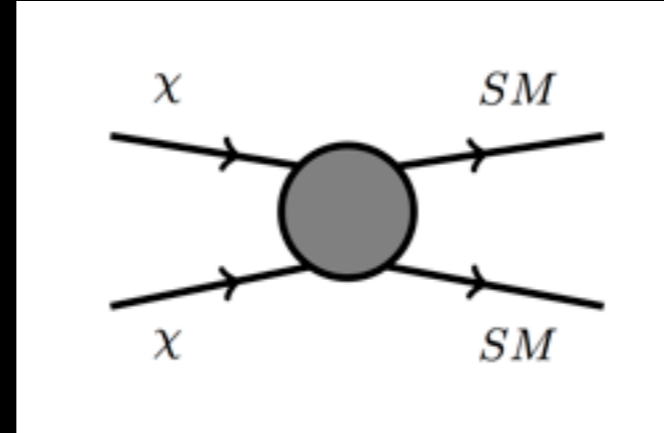
How constrained are these?

Traditional Search Program

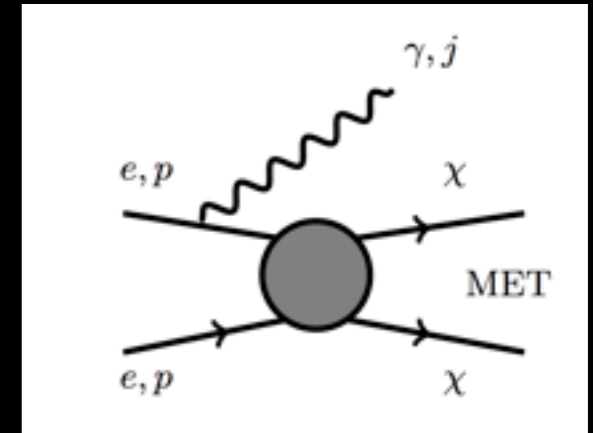
Direct Detection



Indirect Detection



Colliders



Insensitive to sub-GeV thermal DM

Existing Coverage*

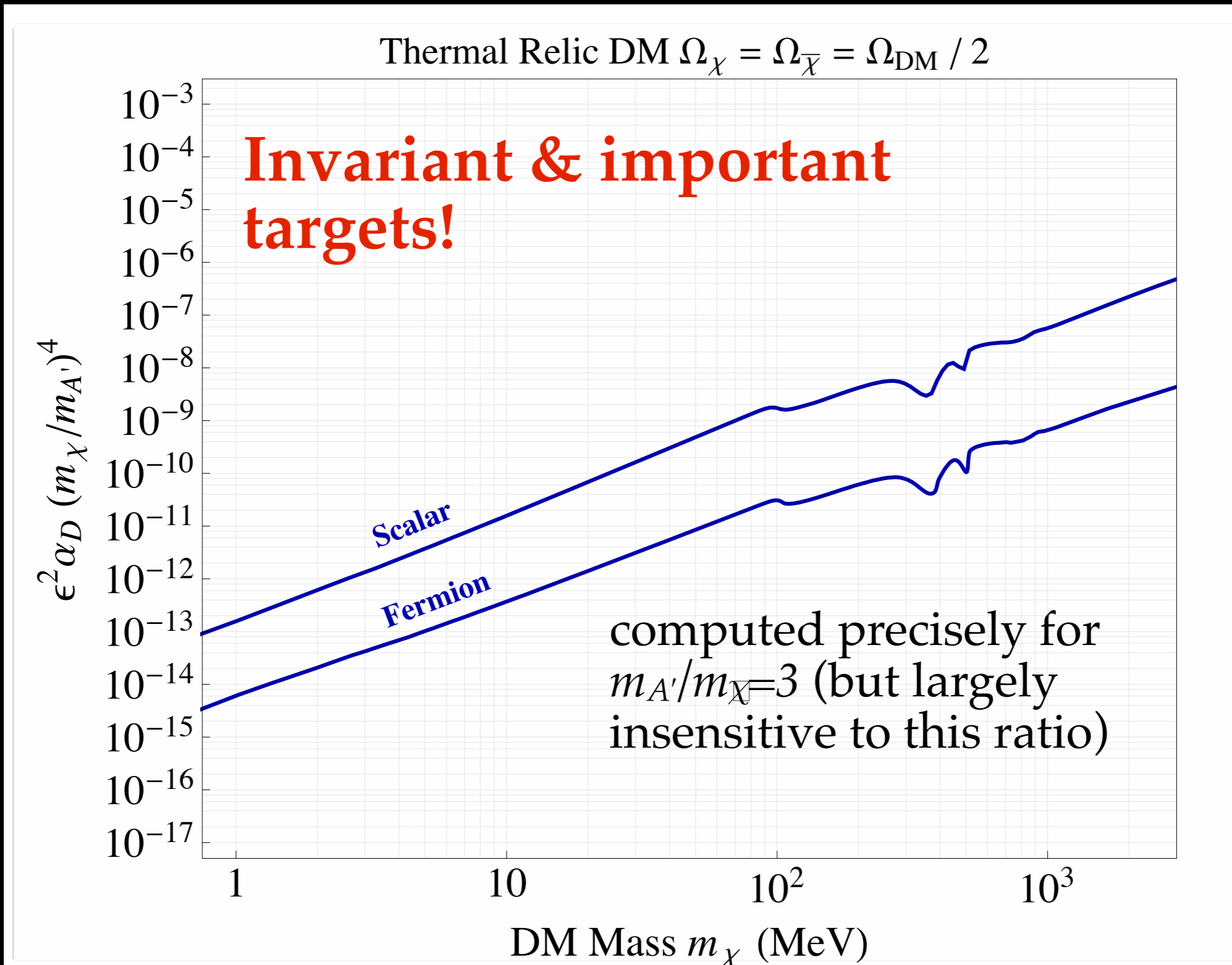


SUSY "WIMP"

Viabale Thermal Models

* = elastic & symmetric

Thermal Target



Bounds: Electron Direct Detection

Xenon10 electron scattering constrains

$$\sigma_{\chi e} \propto \epsilon^2 \alpha \alpha_D \frac{\mu_{\chi e}^2}{m_{A'}^4}$$

Theorist reinterpretation
Essig Mardon, Papucci, Volansky, Zhong
1206.2644

Same scaling as thermal relic target, compare
by plotting:

$$y \equiv \epsilon^2 \alpha_D \left(\frac{m_\chi}{m_{A'}} \right)^4 \quad \text{vs.} \quad m_\chi$$

Model Dependent!

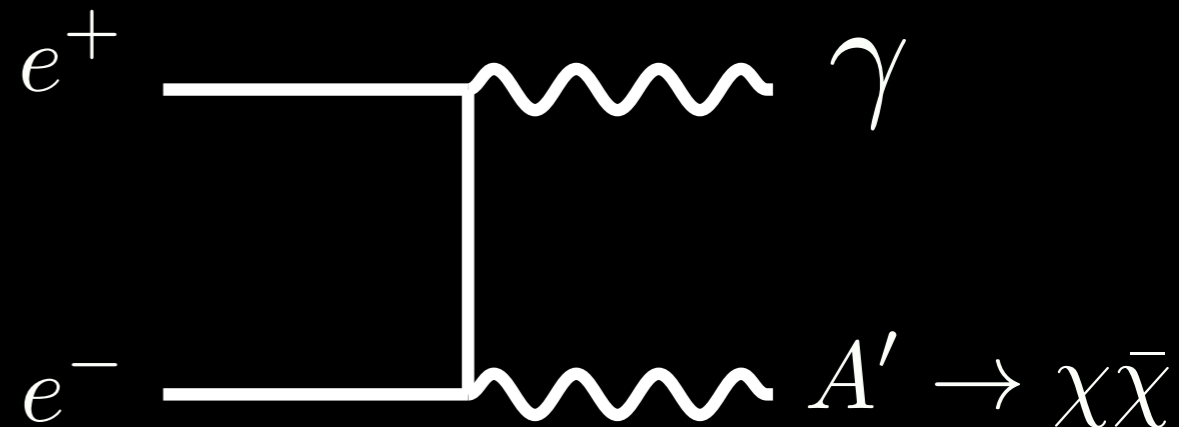
Trivially evaded with inelastic coupling

Bounds: Colliders/B-Factories

BaBar Production:

$$\sigma \sim \frac{\epsilon^2 \alpha}{E_{CM}^2}$$

$$\sim y \times \frac{1}{g_D^2} \left(\frac{m_{A'}}{m_\chi} \right)^4$$



Theorist reinterpretation

Essig, Mardon, Volansky 1309.5084

Izaguirre, GK, Schuster, Toro 1307.6554

Not intuitive: must assume some dark coupling (and mass ratio) to compare with thermal target

Conservative to choose large DM coupling (order-one mass ratio) demands tiny visible coupling for fixed annihilation rate

Bounds: CMB

If* DM can annihilate during CMB epoch, it can reionize Hydrogen

$$\sigma v \propto \epsilon^2 \alpha \alpha_D \left(\frac{m_\chi}{m_{A'}} \right)^4 \frac{1}{m_\chi^2}$$

Slatyer et. al. 1206.2644
Theorist reinterpretation

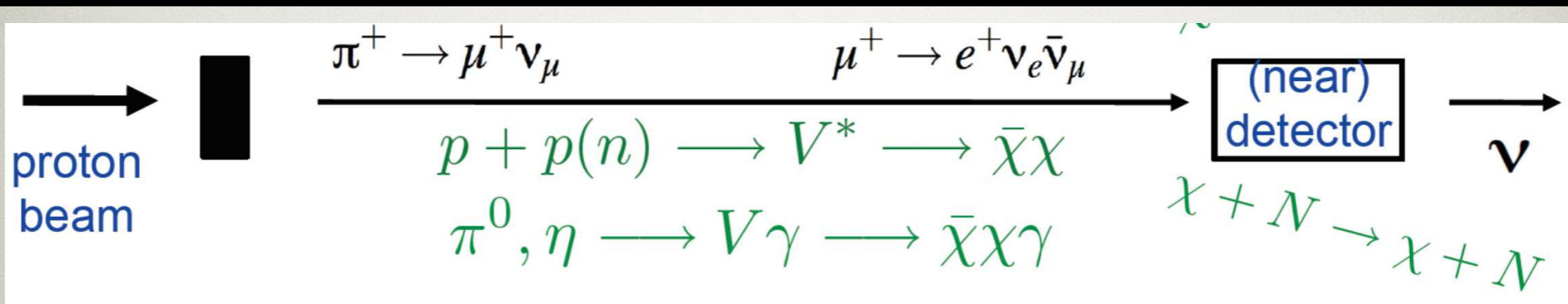
Same scaling as thermal relic target, compare by plotting:

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Model Dependent!

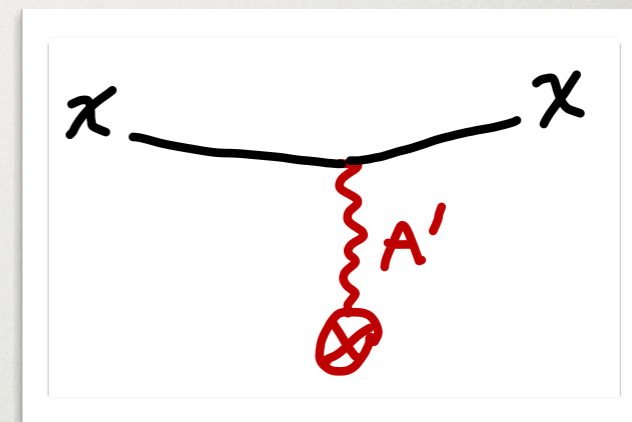
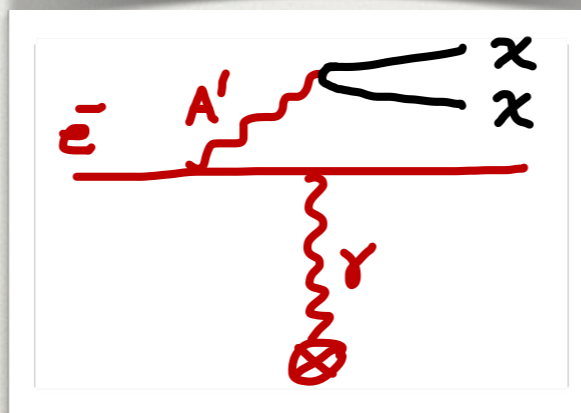
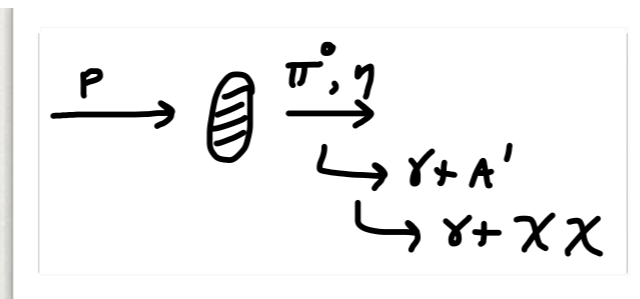
Weaker [absent] for p-wave [inelastic] coupling!

Bounds: Beam Dumps



0906.5614,
1107.4580,1205.3499
Batell, DeNiverville,
McKeen, Pospelov, Ritz

Izaguirre, Krnjaic, PS
& Toro
PRD.88.114015 and
1403.6826



nuclear dissociation;
nucleon, nucleus, or
electron recoil

Good sensitivity (from theorists) for LSND

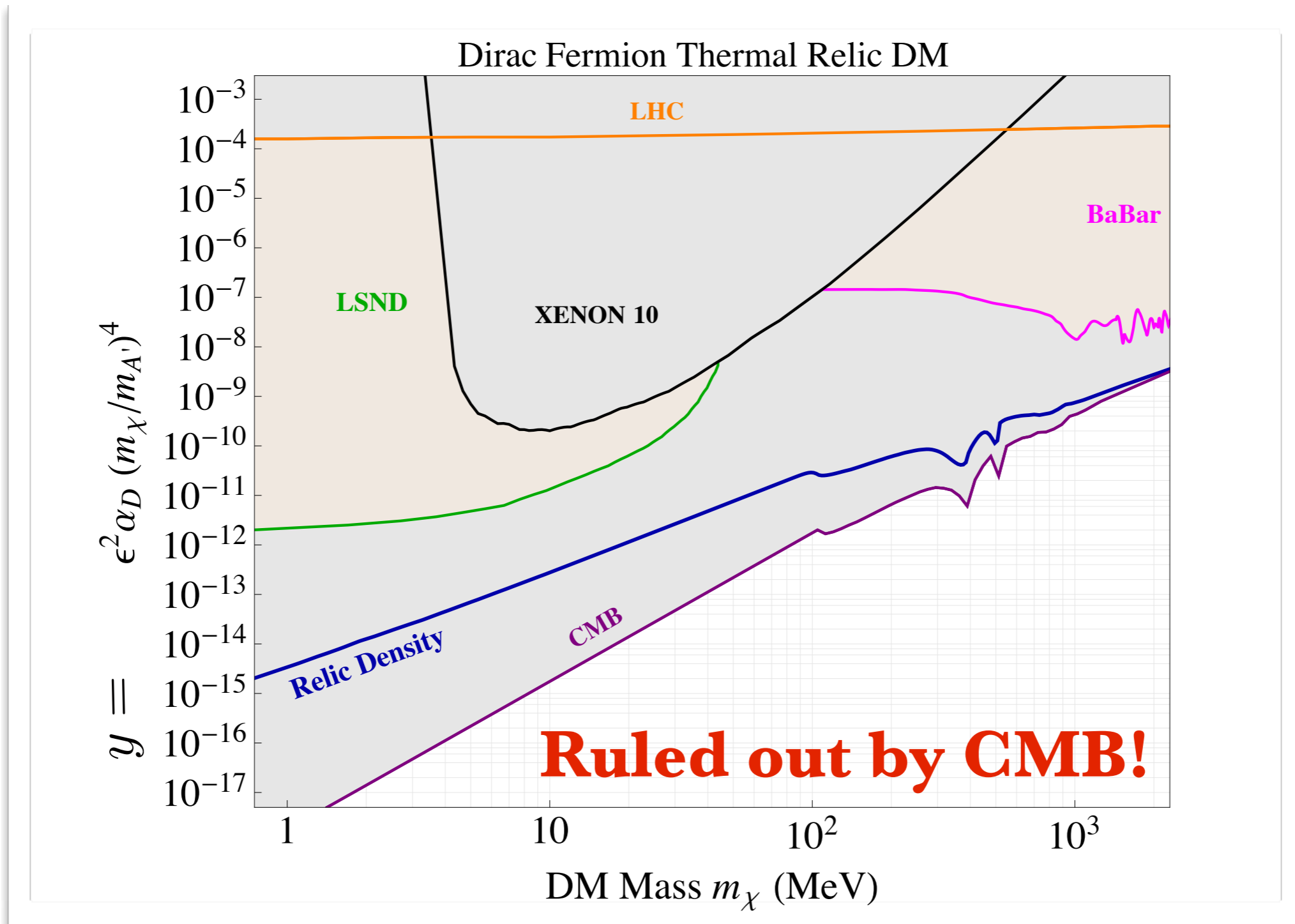
Essig, Batell, Surujon 1406.2698

$$\text{Rate} \propto \epsilon^4 \alpha_D / m_{A'}^2$$

**Conservative
Comparison**

$$\alpha_D, (m_{A'} / m_\chi) \sim \mathcal{O}(1)$$

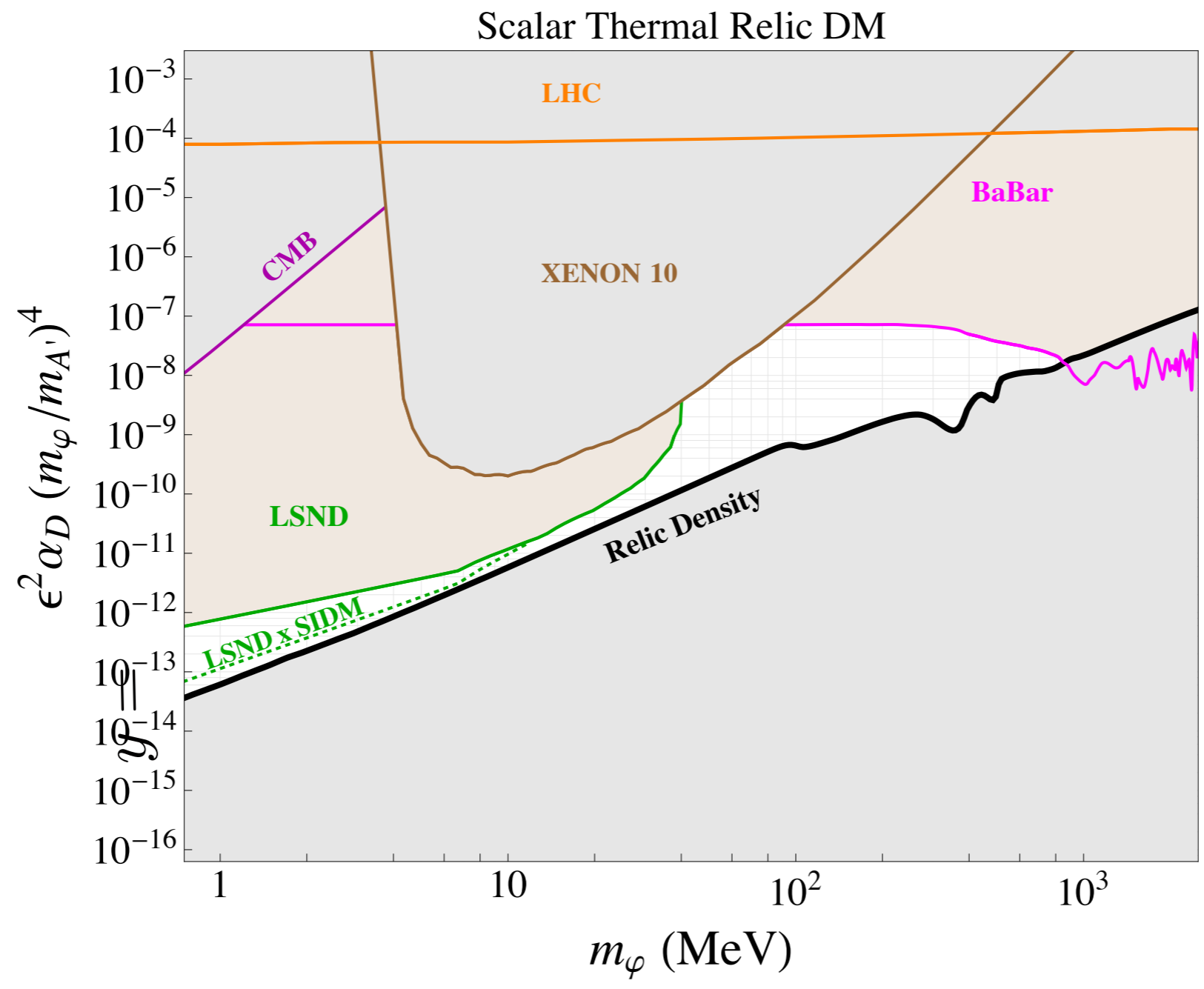
Status: Fermion Symmetric Elastic



BaBar, LSND, LHC: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}}\right)^4 = \frac{1}{81}$

Izaguirre, GK, Schuster, Toro | 504.000 | I

Status: Scalar Symmetric Elastic

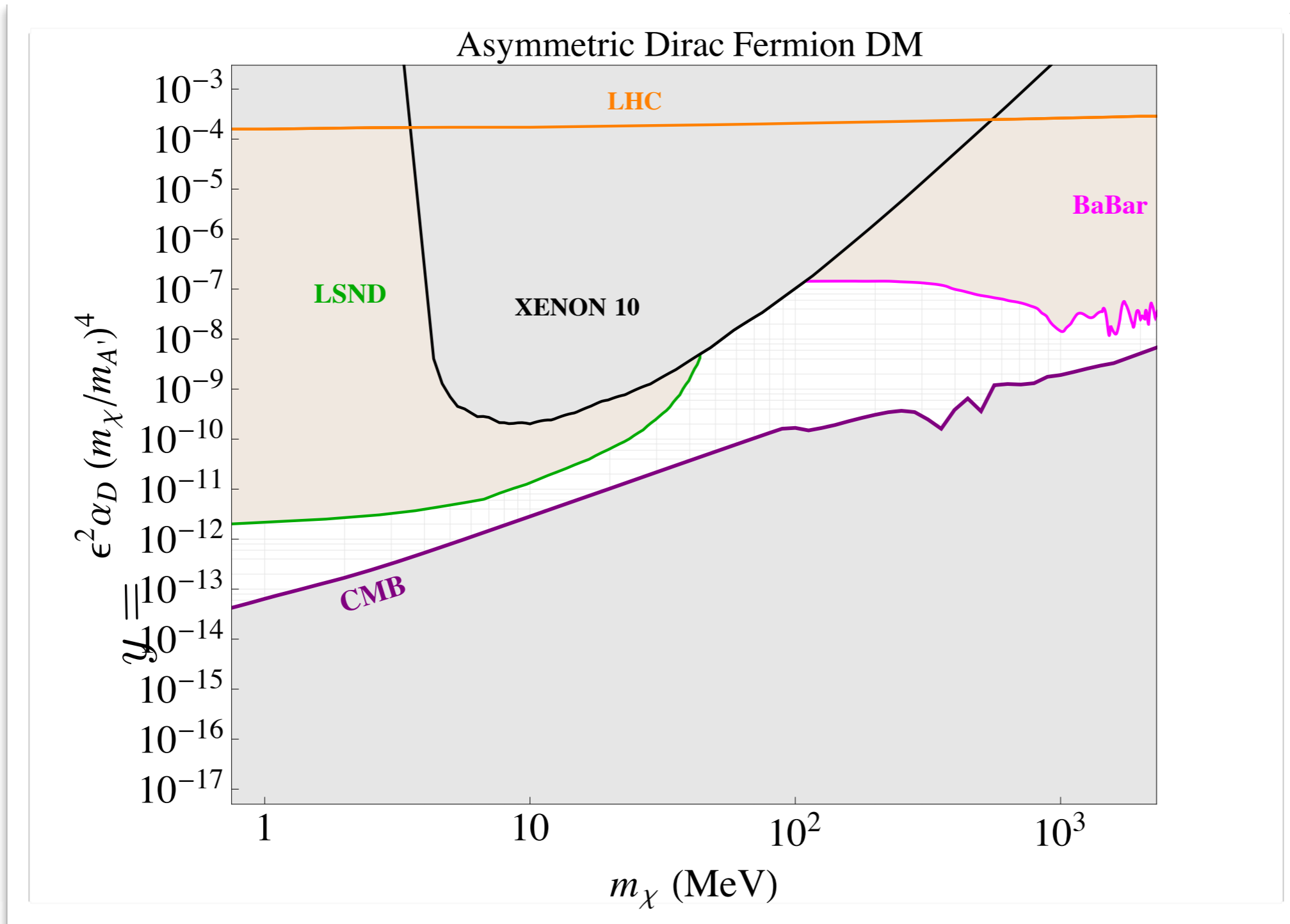


BaBar, LSND, LHC: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}} \right)^4 = \frac{1}{81}$

Izaguirre, GK, Schuster, Toro 1504.00011

Status: Fermion Asymmetric Elastic

$$n_{DM} \neq n_{\overline{DM}}$$

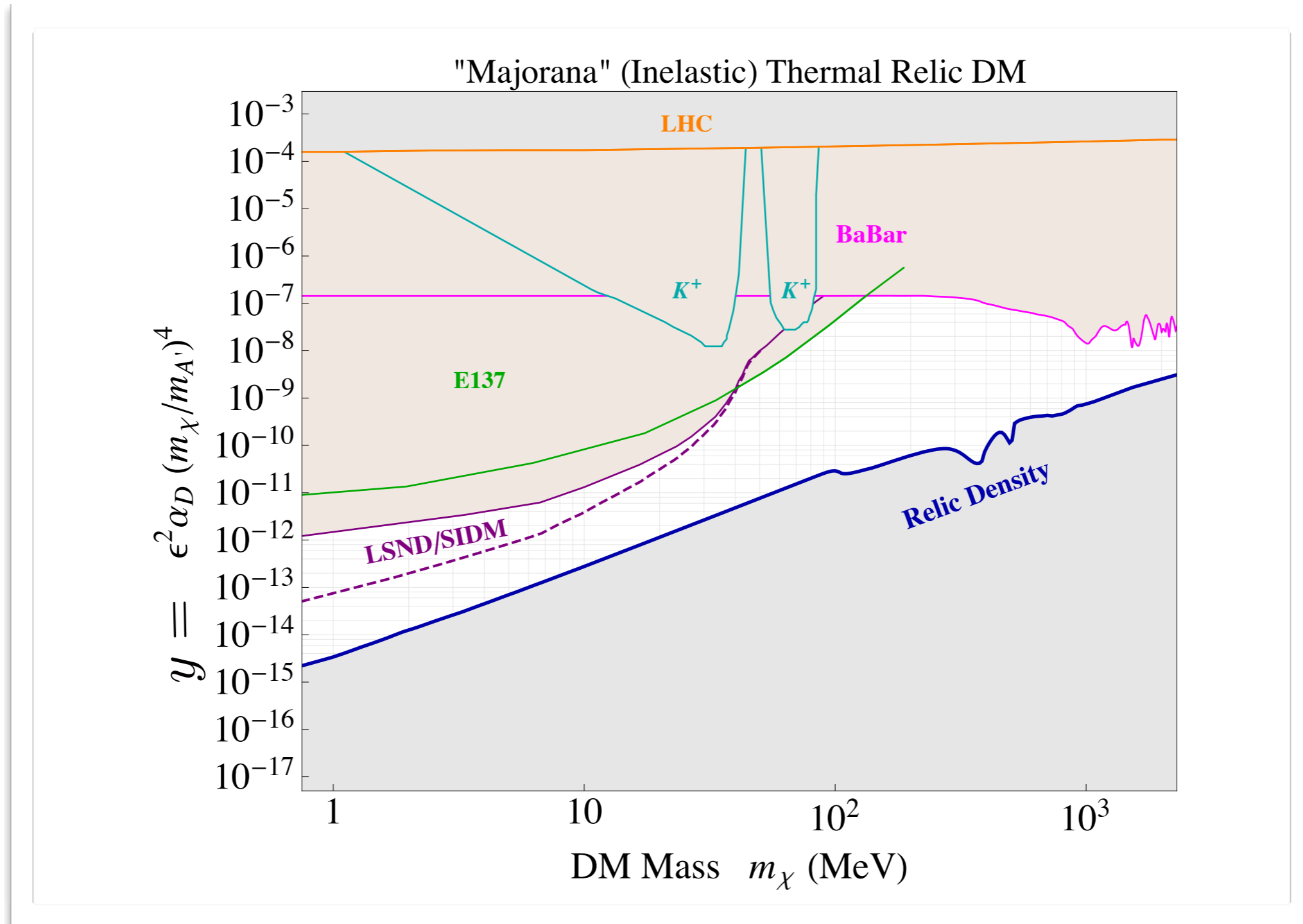
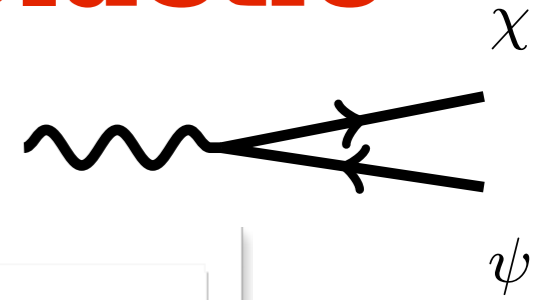


BaBar, LSND, LHC: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}}\right)^4 = \frac{1}{81}$

Izaguirre, GK, Schuster, Toro 1504.00011

Status: Fermion Symmetric Inelastic

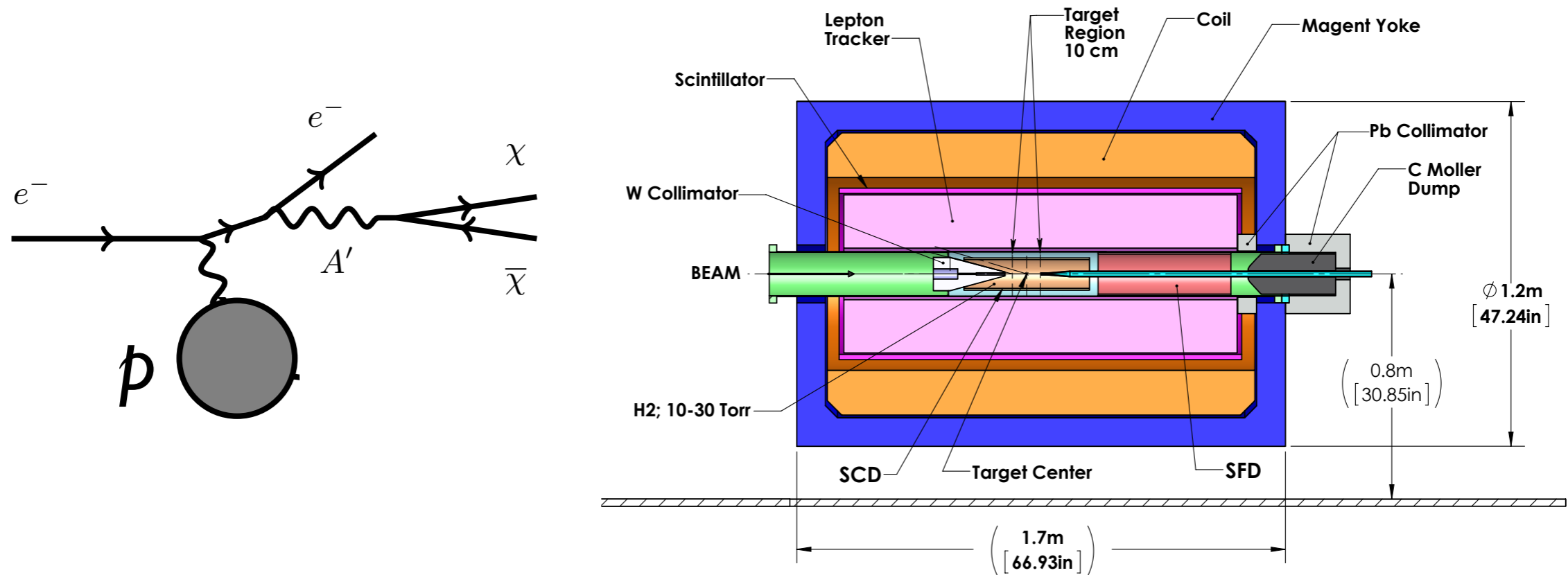
$$\text{keV} < \Delta \ll m_\chi$$



BaBar, LSND, LHC: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}}\right)^4 = \frac{1}{81}$

Izaguirre, GK, Schuster, Toro 1504.00011

DarkLight @ Cornell linac?



Reconstruct e & p kinematics, look for resonance

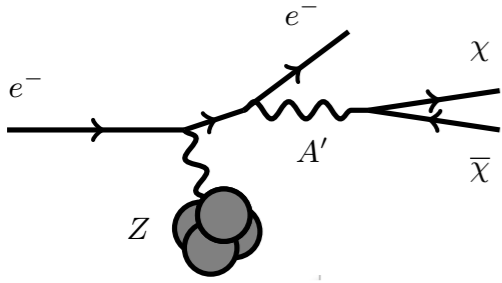
Agnostic: production/detection insensitive to α_D

Good photon rejection required

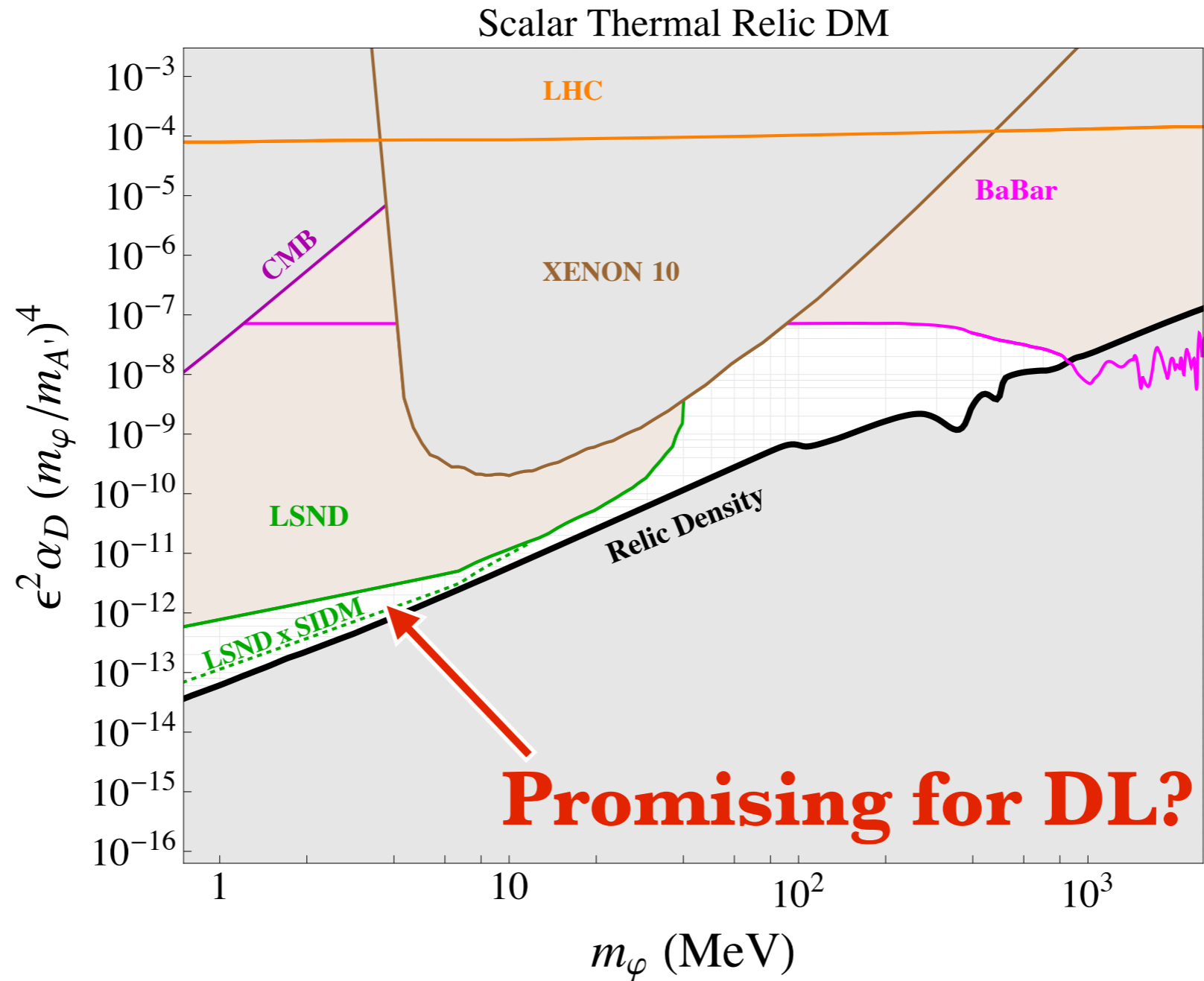
Relevant e-linac features comparable to JLab effort

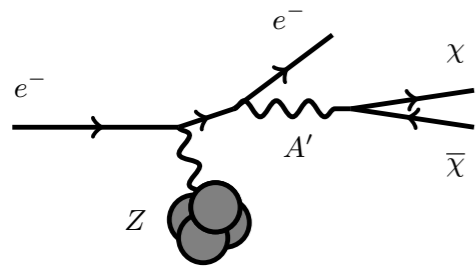
Invisible decay projection (@ JLab)
Kahn, Thaler 1209.0777

Balewski, Fisher et. al. (DarkLight Collaboration)
<http://dmtpc.mit.edu/DarkLight/>

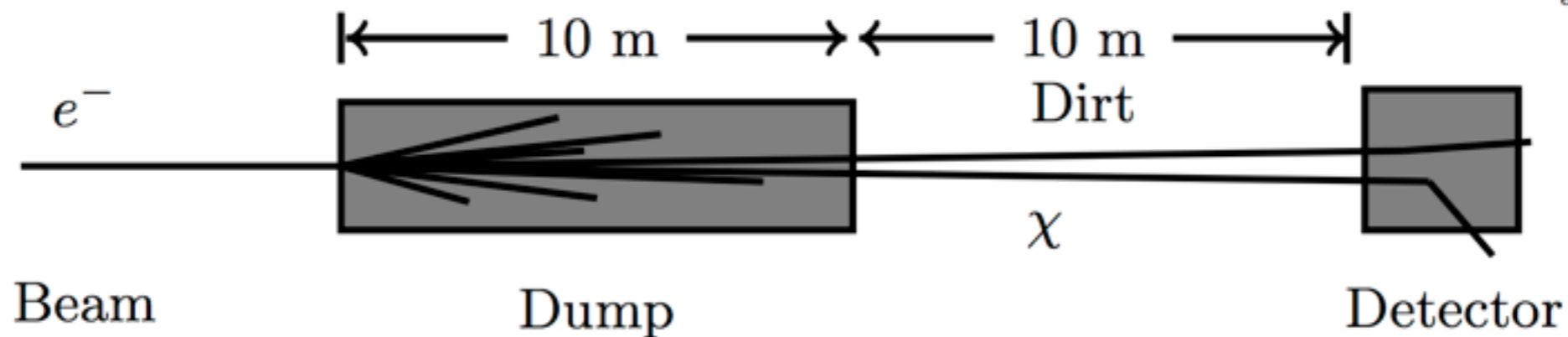
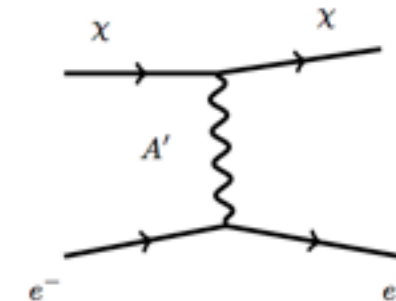


DarkLight @ Cornell linac?





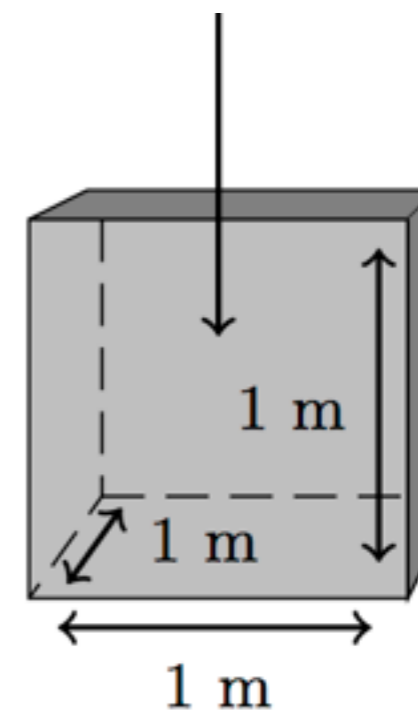
BDX @ Cornell linac?



**Radiative DM production,
Requires downstream scattering**

Beam BG small/negligible

**Cosmic BG reduction necessary
(pulsed beam ok)**

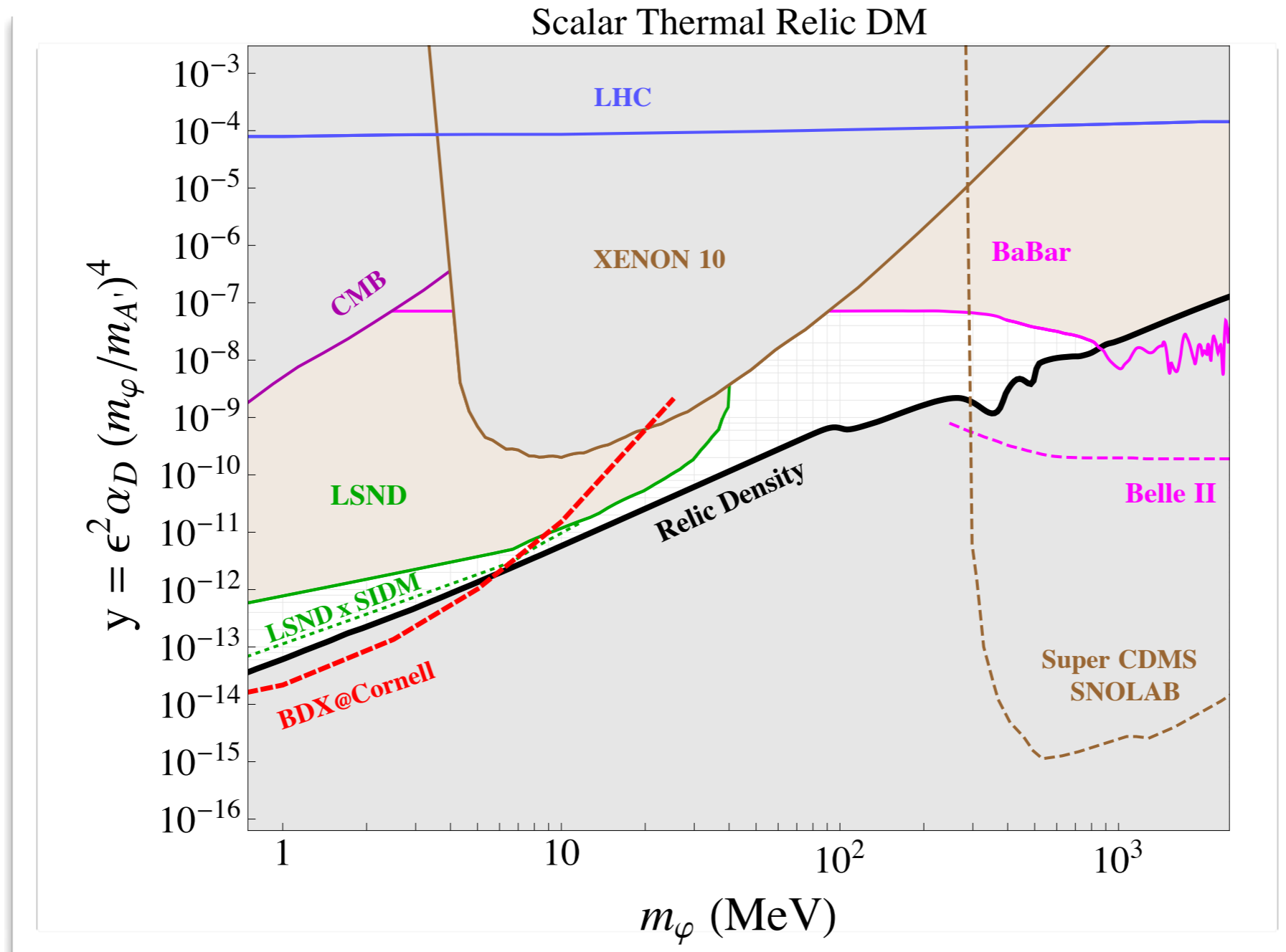


cubic-meter
fiducial

Izaguirre, GK, Schuster, Toro
1307.6554 & 1403.6836

Battaglieri et. al. (BDX collaboration)
1406.3028

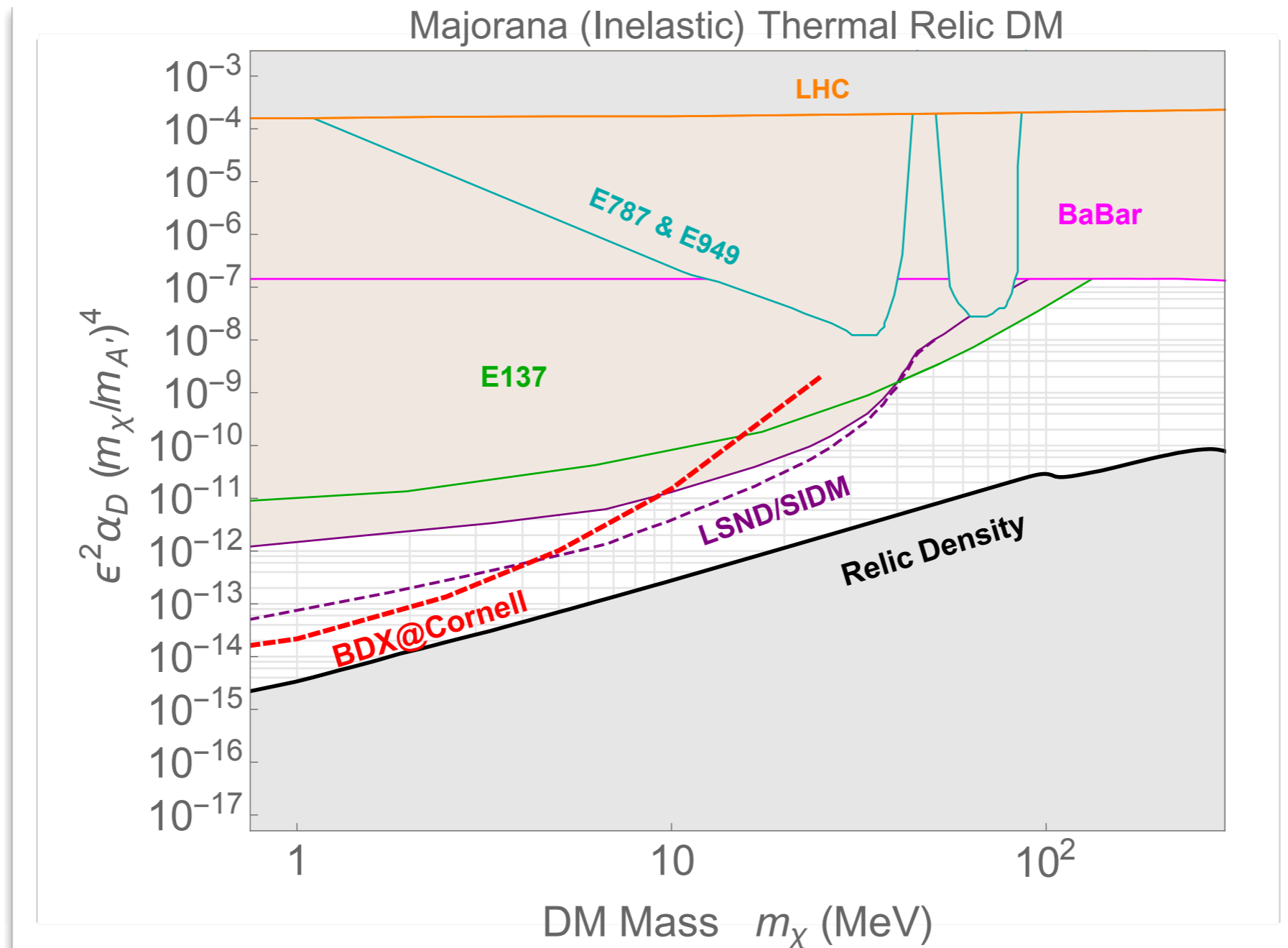
BDX @ Cornell linac?



E(beam) = 400 MeV, 10^{22} EOT, E(recoil) > 50 MeV

BaBar, LSND, LHC, BDX: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}} \right)^4 = \frac{1}{81}$

BDX @ Cornell linac?



E(beam) = 400 MeV, 10^{22} EOT, E(recoil) > 50 MeV

BaBar, LSND, LHC, BDX: $\alpha_D \times \left(\frac{m_\chi}{m_{A'}}\right)^4 = \frac{1}{81}$

Status: Thermal Light DM

Interesting Times Ahead

Fixed Targets

B-factories

Existing Coverage*



SUSY "WIMP"

Viabale Thermal Models

* = elastic & symmetric

Status: Thermal Light DM

Best limits from nontraditional searches

Fixed targets & B-factories especially powerful

Best hope for probing harder scenarios

Most from theorists reinterpreting old experiments

There are no dedicated efforts (yet!)

contrast w/ billion \$ heavy DM program

Lots of viable param-space left for $< \text{GeV}$ DM

Lots of room for improvement

New way forward @ Cornell?

Thank You